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Article

Accepted Version

Bulgin, C. E. ORCID: https://orcid.org/0000-0003-4368-7386, Embury, O. ORCID: https://orcid.org/0000-0002-1661-7828, Corlett, G. and Merchant, C. J. ORCID: https://orcid.org/0000-0003-4687-9850 (2016) Independent uncertainty estimates for coefficient based sea surface temperature retrieval from the Along-Track Scanning Radiometer instruments. Remote Sensing of Environment, 178. pp. 213-222. ISSN 0034-4257 doi: 10.1016/j.rse.2016.02.022 Available at https://centaur.reading.ac.uk/57628/

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To link to this article DOI: http://dx.doi.org/10.1016/j.rse.2016.02.022

Publisher: Elsevier

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Independent uncertainty estimates for coefficient based sea surface temperature retrieval from the Along-Track Scanning Radiometer instruments

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Abstract

We establish a methodology for calculating uncertainties in sea surface temperature estimates from coefficient based satellite retrievals. The uncertainty estimates are derived independently of in-situ data. This enables validation of both the retrieved SSTs and their uncertainty estimate using in-situ data records. The total uncertainty budget is comprised of a number of components, arising from uncorrelated (eg. noise), locally systematic (eg. atmospheric), large scale systematic and sampling effects (for gridded products). The importance of distinguishing these components arises in propagating uncertainty across spatio-temporal scales. We apply the method to SST data retrieved from the Advanced Along Track Scanning Radiometer (AATSR) and validate the results for two different SST retrieval algorithms, both at a per pixel level and for gridded data. We find good agreement between our estimated uncertainties and validation data. This approach to calculating uncertainties in SST retrievals has a wider application to data from other instruments and retrieval of other geophysical variables. Keywords:

Preprint submitted to Remote Sensing of Environment

March 3, 2016

Sea Surface Temperature, Uncertainty Budget, Remote Sensing, Climate Change Initiative

1 1. Introduction

Uncertainty is inherent in all geophysical measurements and must be ap-2 propriately characterised for their scientific application. Data providers have 3 a responsibility to communicate the levels of uncertainties associated with 4 their products and inform data users of the correct methodology for using 5 uncertainty information provided. Within the Sea Surface Temperature Cli-6 mate Change Initiative (SST CCI) project (Hollmann et al., 2013; Merchant et al., 2014) we aim to provide an uncertainty budget for every SST value 8 provided in products (skin temperature, SST at 0.2 m depth and spatially 9 averaged SST). We aim to derive uncertainty estimates independently of SST 10 validation datasets, allowing validation of both the SST values and their as-11 sociated uncertainty. 12

The terms 'error' and 'uncertainty' are sometimes used interchangeably, 13 but have distinct standard definitions that will be adhered to throughout this 14 paper. Error is the difference between a measured value and the true value of 15 the measurand (JCGM, 2008; Kennedy, 2014). In practice we know neither 16 the true value nor therefore the error for a particular measurement. However 17 the distribution of the errors can often be estimated and this distribution 18 characterises the uncertainty in the measured value. Formally, uncertainty 19 is a parameter characterising the dispersion of values that could reasonably 20 be attributed to the measured value (JCGM, 2008). To quantify uncertainty 21 in this paper we quote one standard deviation of the error distribution. 22

It is common to provide generic uncertainty estimates for remotely sensed 23 SST derived via comparison with in-situ datasets during validation activites. 24 The standards of the Group for High Resolution Sea Surface Temperature 25 (GHRSST) specify the provision in all datasets of single sensor error statis-26 tics (SSES). For pragmatic reasons, SSES are defined to comprise the mean 27 difference and standard deviation of remotely sensed SST matched to a 'refer-28 ence' dataset (GHRSST Science Team, 2010). Drifting buoy SSTs are often 29 used as the 'reference'. Mean and standard deviation validation statistics 30 are often provided as globally invariant dataset specific values (May et al., 31 1997; Reynolds et al., 2002; Casey and Cornillon, 1998). An additional field 32 indicating the retrieval quality level can be specified at pixel resolution pro-33 viding information on the likelihood of cloud contamination, noise lamplifi-34 cation at extreme satellite zenith angles or input data quality (Donlon et al., 35 2007; Kilpatrick et al., 2001). An extension of this approach is the MOD-36 erate Resolution Infrared Spectrometer (MODIS) algorithm, which provides 37 validation-based uncertainty information stratified by season, latitude, sur-38 face temperature, satellite zenith angle, a selected brightness temperature 39 difference, SST quality level and day/night (Castro et al., 2010). 40

Sources of uncertainty in remotely sensed SST are intrinsic to the retrieval process and the data utilised. Uncertainties vary from pixel to pixel due to local changes in instrument noise, satellite viewing geometry and atmospheric conditions. We present here a method of estimating SST retrieval uncertainty that accounts for these factors at the pixel level. There are a number of sources of uncertainty in SST measurement and the need to differentiate the effects of random, and systematic errors has been previously noted (Reynolds et al., 2002; Casey and Cornillon, 1998; Merchant et al., 2012; Kennedy,
2014). Gridding of products introduces sampling uncertainties and a number
of studies have considered these when constructing global or regional SST
datasets from in-situ observations (She et al., 2007; Folland et al., 2001;
Rayner et al., 2006; Morrissey and Greene, 2009; Jones et al., 1997; Brohan
et al., 2006).

In this paper, we consider uncorrelated and locally systematic effects con-54 tributing to SST uncertainty. The random or uncorrelated effects arise from 55 noise in the satellite brightness temperature, which propagates into the re-56 trieved SST. Locally systematic effects cause errors that are correlated on 57 synoptic scales of atmospheric variability and are related to the retrieval 58 method itself interacting with changes in atmospheric properties (Minnett, 59 1991; Barton, 1998; Le Borgne et al., 2011; Minnett and Corlett, 2012; 60 Embury and Merchant, 2012; Merchant et al., 2012). We also discuss un-61 certainties from large scale systematic effects (spatially coherent on larger 62 scales than synoptic features). In a companion paper (Bulgin et al., 2016) 63 we derive a method for calculating sampling uncertainty in gridded products 64 due to incomplete sampling of observations in each grid cell, primarily as a 65 result of cloud cover. In this paper, we use reuslts from Bulgin et al. (2016), 66 and, for completeness, show how sampling uncertainty combines with other 67 components of uncertainty in gridded products. 68

The remainder of the paper is structured as follows. Section 2 describes the theory behind the calculation of uncertainties, their propagation and how this is applied to different levels of SST data (orbit data and gridded products). Section 3 describes how an initial uncertainty budget is constructed ⁷³ from errors originating from random, locally correlated and sampling effects.
⁷⁴ In Section 4 we present a validation of our uncertainty budget and in Section
⁷⁵ 5 provide a discussion of the results. We conclude the paper in Section 6.

⁷⁶ 2. Uncertainty Calculation and Propagation

⁷⁷ We construct an uncertainty budget for SST measurements in CCI prod-⁷⁸ ucts comprised of uncertainty components arising from random, locally sys-⁷⁹ tematic, large-scale systematic and sampling effects. The full equation for ⁸⁰ the propagation of uncertainty in a variable y, (u(y)), given that y is related ⁸¹ to input quantities x_i via $y = f(x_1, ..., x_n)$, is defined as equation (1) in the ⁸² Guide to the Expression of Uncertainty in Measurement (GUM) (JCGM, ⁸³ 2008).

$$u^{2} = \sum_{i}^{n} \left(\frac{\partial f}{\partial x_{i}}\right)^{2} u_{i}^{2}(x_{i}) + 2\sum_{i=1}^{n-1} \sum_{j=i+1}^{n} \left(\frac{\partial f}{\partial x_{i}}\right) \left(\frac{\partial f}{\partial x_{j}}\right) u(x_{i}, x_{j})$$
(1)

Uncertainty is expressed with respect to (y) in the GUM, and we repro-84 duce this notation throughout the paper. However, in Earth Observation, 85 we conventionally relate a retrieval estimate \hat{x} to observations y ie. $\hat{x} = f(y)$ 86 which is the reverse convention. The first term in equation (1) describes the 87 propagation of uncertainties from uncorrelated errors. These can be added 88 in quadrature with the differential term $(\partial f/\partial x_i)$ defining the sensitivity of 89 the total uncertainty to each uncertainty component. The second term de-90 scribes the propagation of uncertainty terms arising from correlated errors. 91 This term sums the uncertainty components from correlated errors for each 92 pair of input variables $(x_i \text{ and } x_j)$ found as the product of the sensitivity for 93

⁹⁴ both x_i and x_j and the covariance between them, $u(x_i, x_j)$. The factor of '2' ⁹⁵ is included, as for each pair, each is equally correlated with the other.

Equation (1) can also be written in the form of equation (2) where the uncertainty is expressed as the sum over all pairs of input variables and the covariance term is expressed as the product of the standard uncertainty in x_i , written u_i , in x_j , written u_j , and of the correlation of errors in x_i and x_j , written r_{ij} .

$$u^{2} = \sum_{i=1}^{n} \sum_{j=1}^{n} \frac{\partial f}{\partial x_{i}} \frac{\partial f}{\partial x_{j}} u_{i} u_{j} r_{ij}$$

$$\tag{2}$$

Equation (2) applies fairly generically to any transformation $y = f(x_i, ..., x_n)$ for which the sensitivity parameters $(\partial f / \partial x_i)$ are adequately constant over the range $x_i - u_i$ to $x_j + u_j$; it is a first order approximation. Because we will use the results later, we illustrate the use of equation (2) for calculating the uncertainty in the mean SST from a number of observations. If $f = \sum_{i=1}^{n} x_i/n$, where each x_i is a contributing SST value, then the sensitivity parameter is $\partial f / \partial x_i = 1/n$ giving:

$$u^{2} = \frac{1}{n^{2}} \sum_{i=1}^{n} \sum_{j=1}^{n} u_{i} u_{j} r_{ij}$$
(3)

We can consider three limiting cases. First assume errors are uncorrelated between pixels. We can then put $r_{ij} = \delta_{ij}$, where $\delta_{ij} = 1$ for i = j, and $\delta_{ij} = 0$ for $i \neq j$. In this case, the uncertainty in the mean is scaled by the familiar $\frac{1}{\sqrt{n}}$, reduction in uncertainty, because

$$u^{2} = \frac{1}{n^{2}} \sum_{i=1}^{n} \sum_{j=1}^{n} u_{i} u_{j} \delta_{ij}$$
(4)

$$=\frac{1}{n^2}\sum_{i}^{n}u_i^2\tag{5}$$

Second, consider the case $r_{ij} = 1$, which means errors fully correlate between contributing SSTs. Equation (3) becomes

$$u^{2} = \frac{1}{n^{2}} \sum_{i=1}^{n} \sum_{i=j}^{n} u_{i} u_{j}$$
(6)

$$=\frac{1}{n^2}\left(\sum_{i=1}^n u_i\right)^2\tag{7}$$

implying $u = \frac{1}{n} \sum_{i=1}^{n} u_i$ ie. the uncertainty is the average uncertainty of the contributing SSTs.

Third, consider the case $r_{ij} = \delta_{ij} + (1 - \delta_{ij})r$ - all SSTs have the same error correlation with other SSTs. Substituting into equation (3) gives

$$u^{2} = \frac{1}{n^{2}} \sum_{i}^{n} \sum_{j}^{n} u_{i} u_{j} [\delta_{ij} + (1 - \delta_{ij})r]$$
(8)

$$= \frac{1}{n^2} \sum_{i=1}^{n} \sum_{j=1}^{n} u_i u_j [r + (1-r)\delta_{ij}]$$
(9)

$$= \frac{r}{n^2} \left(\sum_{i=1}^n u_i\right)^2 + \frac{(1-r)}{n^2} \left(\sum_{i=1}^n u_i^2\right)$$
(10)

This form yields the previous results as special cases (r = 0 and r =119 1). Constant r_{ij} for $i \neq j$ is in practice unlikely to be exact for a real 120 situation, but may be a useful approximation in some cases, avoiding the 121 need to estimate r_{ij} for every contributing pair.

122 3. Uncertainty Budget Components

123 3.1. Uncorrelated Effects

Random errors in SST estimation from satellite data arise from noise 124 in the satellite observations. The signal recorded by a typical radiometer 125 is a voltage measured across a detector, digitised and recorded as counts. 126 In the operational calibration, a linear radiance is calculated in the form 127 $radiance = gain \times counts + offset$ where the gain and count parameters are 128 calculated during instrument calibration (Smith et al., 2012). A non-linearity 129 adjustment is then applied to the longwave channels (Smith et al., 2012) for 130 which the associated uncertainty has not been calculated. In this analysis 131 we simply take the detector noise in the measured counts and propagate this 132 into our geophysical retrieval. In a coefficient based retrieval, SST is calcu-133 lated from a pre-defined linear or nearly linear (Anding and Kauth (1970); 134 Deschamps and Phulpin (1980); Kilpatrick et al. (2001); May et al. (1997); 135 McMillan and Crosby (1984), and further references within Merchant (2013)) 136 combination of the observed brightness temperatures. Brightness tempera-137 ture uncertainty is characterised using channel-specific noise equivalent dif-138 ferential temperature (NEdT). This uncertainty is then propagated into the 130 SST retrieval uncertainty. 140

We illustrate the propagation of errors from random effects using data from the polar orbiting Advanced Along Track Scanning Radiometer (AATSR) aboard the Envisat satellite. Envisat was in a sun synchronous orbit with an equator overpass time of 10.00 am. AATSR made observations in seven spectral bands covering the visible and infrared spectrum at two viewing geometries: nadir $(0 - 22^{\circ})$ and forward $(52 - 55^{\circ})$. SST can be derived using the nadir infrared channels only, or using both the nadir and forward views. We consider here the propagation of uncertainties through two different retrievals: 'N2' using the 11 and 12 μ m channels in the nadir view only and 'D2' using the 11 and 12 μ m channels in both views. The formula used here for estimating coefficient based SSTs from satellite data is:

$$\hat{x}_{SST} = a_0 + \sum_k a_k y_k \tag{11}$$

¹⁵² Where y_k refers to each channel used in the retrieval, a_0 is an offset and a_k ¹⁵³ are channel specific coefficients. Note that here $\hat{x} = f(y)$, in contrast to usage ¹⁵⁴ in Section 2 (as previously noted). These coefficients vary with the context in ¹⁵⁵ which the observation is made, according to the viewing geometry and total ¹⁵⁶ column water vapour (TCWV), but are predefined. The error (difference ¹⁵⁷ between the measured value and true value) for a given SST can be defined ¹⁵⁸ as:

$$e_{SST} = \sum_{k} a_k e_{y_k} \tag{12}$$

This is a linear combination of the errors in the brightness temperatures 159 in each chanel (denoted by 'k') multiplied by the coefficient used in the 160 retrieval. In practice, we do not know the true SST value nor therefore the 161 error on each individual measurement, but we can simulate a 'typical' error 162 field from our knowledge of the NEdT in each channel. We illustrate this in 163 panels 1 and 2 of Figure 1 which show simulated error fields for the nadir 164 view of the 11 and 12 μ m channel at pixel resolution (1 km at nadir for 165 AATSR). These are constructed using a Gaussian random number generator 166 selecting values from a distribution with 0.0 °C mean and 0.05 °C standard 167

Retrieval	Channel	Sec(Sat Zenith Angle)	a_1 coefficient
N2	$11 \ \mu { m m}$	1.0	2.04314
N2	$12 \ \mu { m m}$	1.0	-1.02542
D2	$11 \ \mu { m m}$	1.0	4.65371
D2	$11 \ \mu { m m}$	1.76	-1.65009
D2	$12 \ \mu \mathrm{m}$	1.0	-3.27043
D2	$12 \ \mu \mathrm{m}$	1.76	1.27186

Table 1: Coefficients for each channel used to calculate SST in the 'N2' and 'D2' retrievals in Figure 1.

deviation representing NEdT estimates for the two channels (Embury and
Merchant, 2012). Errors vary in magnitude from pixel to pixel and can be
either positive or negative in sign.

Panels 3 and 4 of Figure 1 show the propagation of these simulated error 171 fields in a N2 and D2 retrieval. For the purpose of this illustration we assume 172 a fixed view angle and TCWV (23 kg m^{-2}) across the image giving coefficients 173 (a_k) dependent only on channel, as shown in Table 1. Under normal retrieval 174 conditions these would vary slightly on a per-pixel basis. The coefficients 175 are specified to five decimal places (Merchant and LeBorgne, 2004). Further 176 discussion of error inherent in the retrieval process is provided in Section 177 3.2. As indicated in equation (12) the uncorrelated errors in a given retrieval 178 are the sum of the errors in each channel, and therefore the total errors are 179 smaller in the N2 retrieval than the D2 retrieval (which uses four channels 180 with generally heavier weights). 181

182

Many users require gridded Level 3 products generated from full reso-

lution data. When generating gridded products, the average SST can be
calculated using the arithmetic mean:

$$\hat{x}_{GriddedSST} = \frac{1}{n} \sum_{i=1}^{n} \hat{x}_{SST(i)}$$
(13)

Where n is the number of observations (i) in the grid cell. The alternative 185 would be to calculate a weighted mean based on the per-pixel uncertainties, 186 but we choose the arithmetic mean as it gives equal weight to all measure-187 ments across the grid cell and therefore represents a mean across the geo-188 physical variability within the grid cell. Panels 5 and 6 show the arithmetic 189 mean of the errors over a 5 x 5 pixel grid cell, approximately representing the 190 creation of 0.05° Level 3 products. The range in the mean error is naturally 191 smaller in the gridded product, but remains larger for the D2 retrieval than 192 the N2 retrieval. 193

In practice, when retrieving SST from satellite observations we don't ex-194 plicitly know the error in either the brightness temperatues or SST. We need, 195 however, to estimate the uncertainty in the SST retrieval. Given estimates 196 of NEdT, this is an example of standard uncertainty propagation. 'Standard 197 uncertainty' is the standard deviation of errors in each channel brightness 198 temperature, estimated to be of the order of 0.05 K for both the 11 and 12 199 μ m channels of AATSR (Embury and Merchant, 2012). The propagation of 200 uncorrelated uncertainty components is shown in equation (5) where uncer-201 tainties are added in quadrature. Applying this to equation (11), in the first 202 instance to give the per pixel uncertainty, and differentiating with respect to 203 each channel (y_k) used in the retrieval gives: 204

$$u_i = \sqrt{\sum_k a_k^2 u_{y_k}^2} \tag{14}$$

For a gridded product using the arithmetic mean, the uncertainty in the mean of the contributing pixels is

$$u_{GriddedSST} = \frac{1}{\sqrt{n}} \sqrt{\frac{\sum u_i^2}{n}}$$
(15)

For fixed coefficients and a constant error in the brightness temperatures 207 (0.05 K) as in Figure 1, there is an invariant uncertainty value for each re-208 trieval algorithm (N2 and D2) at the pixel level. When creating a real SST 209 product, NEdT varies as a function of both channel and brightness temper-210 ature. For N2 retrievals in the example provided, this invariant uncertainty 211 value is 0.11 K and for D2 retrievals 0.25 K. Uncertainties in gridded aver-212 ages reduce by $\frac{1}{\sqrt{n}}$ giving uncertainty estimates of 0.02 K and 0.05 K for N2 213 and D2 retrievals over fully observed grid cells. In practice, many grid cells 214 in Level 3 products are not fully observed due to cloud cover. This reduces 215 the number (n) of observations available and increases the uncertainties from 216 random effects. This is illustrated in panels 7 and 8 of Figure 1 for N2 and 217 D2 retreivals. A cloud mask has been superimposed on the simulated data at 218 the per-pixel level and uncertainties propagated into the 5x5 pixel product. 219 Observing only part of a given grid cell additionally introduces sampling un-220 certainty, discussed briefly in Section 3.4 and more fully in the companion 221 paper (Bulgin et al., 2016). 222

223 3.2. Locally Systematic Effects

Uncertainties from locally systematic effects arise from ambiguities in or 224 limitations of the SST retrieval algorithm. Coefficient based retrievals for 225 the ATSR instruments in Phase 2 of the SST CCI will use coefficients from 226 the ATSR Reprocessing for Climate (ARC) project. These are calculated 227 based on radiative transfer simulations which cover a comprehensive range of 228 surface and atmospheric conditions (Embury and Merchant, 2012; Embury 220 et al, 2012). Locally systematic effects therefore vary on synoptic scales 230 consistent with changes in atmospheric conditions. 231

We can characterise the uncertainties arising from locally systematic ef-232 fects in the retrieval scheme using simulation studies. To do this, we take a 233 'true' SST field from Numerical Weather Prediction (NWP) data and simu-234 late the associated brightness temperatures globally as would be observed by 235 the AATSR instrument using the RTTOV radiative transfer model. We can 236 then use these simulated brightness temperatures as input into our retrieval 237 scheme, comparing our retrieved SST with the 'true' SST eg. (Merchant 238 et al., 2009). For any given scene, we can plot the retrieval error field using 239 this methodology as shown in Figure 2. The contour lines in the top pan-240 els show atmospheric pressure and in the bottom two panels TCWV with 241 the spatial distribution of the error field consistent with synoptic scales of 242 pressure in hPa and total column wate vapour (TCWV) in kqm^2 variability. 243 However, features in the SST error field are not simply linked to TCWV 244 distributions, since we see that a single contour line can run through re-245 gions of both positive and negative errors. The ARC retrieval coefficients are 246 banded by TCWV and the observed errors are not simply a bias that can 247

²⁴⁸ be removed from the retrieval. Uncertainty arising from these error effects ²⁴⁹ is characterised in the retrieval as a function of TCWV consistent with the ²⁵⁰ coefficient banding. Panels in the left and right in Figure 2 show the SST ²⁵¹ retrieval error fields for different days, which vary in time as well as space on ²⁵² synoptic scales.

Within the retrieval scheme, uncertainties are calculated as the standard 253 deviation of the error distributions from the simulated data, taking the dif-254 ferences between the 'true' and retrieved SSTs. This is the fitting error of 255 the regression when the coefficients are applied to the simulated data used 256 to generate the coefficients. Figure 3 shows the uncertainties as a function of 257 TCWV for retrievals using different channel combinations at different view-258 ing geometries. For the N2 retrieval using two channels (11 and 12 μ m) the 259 uncertainties increase as a function of TCWV, flattening at higher TCWV's 260 above 45 kg m⁻². With the addition of information from multiple viewing 261 angles $(0-22^{\circ} \text{ and } 52-55^{\circ})$ locally systematic uncertainties are significantly 262 reduced to ~ 0.1 K or lower. 263

Figure 3 also shows the uncertainty from uncorrelated effects as a func-264 tion of TCWV for different channel combinations. Comparing single-view 265 retrieval uncertainties with dual-view uncertainties, the dual-view capability 266 reduces the systematic uncertainty at the expense of the increased retrieval 267 noise. Uncertainties from uncorrelated effects are dependent on both the 268 NEdT for a given channel combination and the coefficients. For the N2 and 269 D2 retrievals large weights are assigned to the 11 and 12 μ m channels which 270 magnifies the uncorrelated uncertainty. ARC coefficients are tuned to assume 271 NEdTs of 0.01 K (smaller than actual values) as they are designed to produce 272

SST products at 0.1° resolution. This has the effect of reducing locally sys-273 tematic uncertainties at the cost of increased uncorrelated uncertainties as 274 these decrease as a function of $1/\sqrt{n}$ when calculating the gridded product. 275 Many SST retrievals also use information from the 3.7 μ m channel at 276 night. The consequence of adding this third channel to the retrieval (results 277 not shown) reduces uncertainty from locally systematic effects to ~ 0.1 K or 278 lower, with larger uncertainties for drier atmospheres. As TCWV increases, 279 the 11 and 12 μ m channels become less sensitive to the surface whilst the 3.7 280 μ m channel remains relatively transparent. SSTs in regions of high TCWV, 281 close to the equator also show less variability which may improve the fit of 282 the retrieval to the training data. For the uncertainties due to uncorrelated 283 effects, including the 3.7 μ m channel in the retrieval results in smaller weights 284 for the 11 and 12 μ m channels reducing the noise amplification. 285

286 3.3. Large Scale Systematic Effects

Other effects can cause SST errors that are correlated on larger scales. For brevity, the uncertainty associated with unknown errors correlated on large scales is hereafter referred to as "systematic uncertainty". (It is taken for granted that any 'known' or 'estimated' systematic errors have been addressed i.e., that any general bias has been quantified and subtracted from data. The systematic uncertainty therefore quantifies the degree of doubt in the measurements associated with what might be termed 'residual biases'.)

All satellite sensors are calibrated prior to launch to a pre-defined standard. The required accuracy for SST measurements from space for climate applications is 0.1 K (Ohring et al, 2005). In some cases the SST algorithm itself is capable of adjusting for some of the systematic errors in calibra-

tion, for example an SST retrieval algorithm that fits regression coefficients 298 to buoys directly will correct for some of the calibration biases as part of 290 the fitting process. This process will also introduce an additional source of 300 uncertainty from unknown errors in the buoy measurement. The buoy data 301 are point measurements at depth whereas the satellite observations are area 302 measurements of skin temperature. If the sensor is poorly characterised this 303 additional uncertainty term can be smaller than the systematic calibration 304 bias. Thermal channels on some sensors seem in practice to have a BT cal-305 ibration accuracy of 0.1 K, judging by the SST accuracy achievable using 306 radiative transfer-based coefficients. 307

The sensor having been calibrated to a certain level, there remain smaller 308 errors, within the specified calibration accuracy, that are unknown. These 309 may vary systematically with scene temperature, general instrument temper-310 ature, the thermal state of the on-board calibration target, the temperature 311 of the detectors, the illumination of the sensor on the space-craft by the Sun, 312 and potentially with many other factors. Sometimes, these effects are suffi-313 ciently evident in flight that they can be diagnosed and corrected for (Cao 314 et al., 2005; Yu et al., 2012; Wang and Cao, 2008; Mittaz and Harris, 2011; 315 Mittaz et al., 2013). There may be a gradual evolution of such systematic 316 calibration effects over time, as the sensor ages, and/or as the platform orbit 317 drifts, changing the illumination and thermal cycling of the sensor. 318

Where satellite datasets are reprocessed, there may be some effort to harmonise the BTs across different sensors. "To harmonise" here means to reconcile the calibration of the observed BTs given the known differences between the sensors; it does not mean that the BTs would be the same for two sensors viewing the same scene; it does mean that the differences would be traceable to known instrumental differences, such as different spectral response functions. The adjustments made to BTs in the light of harmonisation have their own associated uncertainty, and this also is likely to be systematic as defined here. Overall, harmonisation is intended to reduce systematic effects, particularly relative errors between sensors.

It is possible in principle to estimate the systematic uncertainty associated 320 with calibration. There are two possible approaches. The first is to exploit 330 the pre-flight calibration information where an analysis of the potential cal-331 ibration errors has been made. Where such information is available in suffi-332 cient detail in the public domain, it can form the basis of an uncertainty bud-333 get. The second is to exploit near-coincident observations in space between 334 different sensors. Having accounted for instrumental characteristics, differ-335 ences in matched observations can be used to adjust a less-well-calibrated 336 sensor to a better-calibrated sensor. These adjustments have a quantifiable 337 statistical uncertainty, which then provides an estimate of the magnitude of 338 the post-correction systematic uncertainty eg. (Goldberg, 2007). 339

In general, however, calibration uncertainty is not well quantified and 340 propagation of such information into the systematic uncertainty in SST has 341 not been undertaken, to our knowledge. Arguably, for SSTs generated opera-342 tionally for use in numerical weather prediction and real-time oceanography. 343 it has not been necessary. However, in the context of developing repro-344 cessed SST datasets for climate applications, it is an area that needs to be 345 developed. Climate data records require justified uncertainty estimates, par-346 ticularly estimates of their multi-decadal stability, which implies a detailed 347

engagement with understanding and propagating uncertainty from systematic effects throughout the record (Minnett and Corlett, 2012). A metrology (science of measurement) of Earth Observation needs to be developed, to bring relevant metrological principles for developing traceable chains of uncertainty to bear in the context of historic satellite missions.

353 3.4. Sampling Uncertainties

Many users of SST data require gridded products with SST specified as a 354 mean value across the space and time represented by the grid cell. Often grid 355 cells are not fully observed, typically in infrared measurements due to cloud 356 cover, but also in the case of corrupted data or problems with the retrieval 357 process. Data may also be removed from the subsample by conservative cloud 358 detection schemes which can mask clear-sky pixels. The mean SST across 359 the observed pixels may differ from the mean SST across all pixels in the 360 grid cell introducing sampling uncertainty. 361

We cannot explicitly calculate the difference between the SST across the full grid cell and the SST in the available subsample within the retrieval as we do not know the SST of the unsampled pixels. We can however model the sampling uncertainty associated with this process using fully clear-sky data extracts, and we do this as a function of the percentage of the total number of pixels available in the subsample and the standard deviation of the SST in the available pixels.

The full details of the derivation of the sampling uncertainty model are provided in the companion paper (Bulgin et al., 2016). Here we provide only a brief overview, for completeness of the discussions in this paper. In Bulgin et al. (2016) we parameterise sampling uncertainty using a cubic function in the form $(ap^3 + bp^2 + cp + d)$ where a, b, c and d are coefficients dependent on the SST standard deviation in the subsample, and p is the percentage of clear-sky pixels within a given grid cell. This model is therefore applicable to any retrieval scheme with data at the same spatial scale provided that the noise contribution to the SST standard deviation has been subtracted.

378 3.5. Other effects contributing to uncertainty

The propagation of the effects of radiometric noise and the analysis of 379 locally systematic uncertainty discussed has assumed the context of normal 380 clear-sky conditions for each SST retrieval. This neglects the fraction of 381 retrievals that will in practice be made under unusual conditions. These are 382 principally retrievals made for pixels whose classification as clear-sky-over-383 seawater is doubtful, but which have nonetheless passed the cloud screening 384 process. At present, we have no method for estimating this in the uncertainty 385 budget. 386

The first case to consider is 'residual' unscreened cloud contamination. 387 Clouds escape detection if they are sufficiently small and low (warm) or suffi-388 ciently optically thin (e.g., some cirrus). In these cases they can nonetheless 389 affect BTs at the level of several tenths of kelvin. The corresponding im-390 pact on SST depends on how different the cloud impacts on BTs are from 391 the impact of increased water vapour in the atmosphere (which the retrieval 392 algorithms are adapted to deal with). The probability of such cases is con-393 sidered to be greater around the edges of areas correctly identified as cloudy. 394 Note that both the distribution of BT modification by cloud-contamination 395 in pixels falsely considered to be clear sky, as well as the frequency of failure 396 to detect are dependent on the cloud screening system. One could envisage 397

that simulation of a representative range of cloudy situations be carried out to generate such information, but to our knowledge, this has not been done. Given these pieces of information, assessment of the contribution to SST uncertainty could be undertaken by error propagation methods similar to those described earlier. At present, however, the contribution of this effect to SST uncertainty is not estimated.

The second case to consider is atmospheric aerosol of a form and optical 404 depth outside the range of circumstances for which the retrieval coefficients 405 are designed. Again, to the degree that the aerosol affects the BTs differently 406 to water vapour [e.g., Merchant et al. (2006)], SST errors may be induced 407 of unknown size. While aerosol events trigger cloud detection if the optical 408 depths are sufficiently great, there is a regime where SST retrievals can be 409 affected, the effect in most cases being to make the retrieved SST too cold. 410 Again, the contribution of this effect to SST uncertainty is not estimated. 411

The third case relates to sea ice being present within the pixel for which SST is retrieved. If the ice is not too cold and is relatively dark (circumstances that often go together in the formation of new ice), the ice may not be detected. Similar considerations apply as to missed residual cloud or aerosol, and this contribution to uncertainty again is not presently estimated.

There are a number of further effects contributing to SST uncertainty that are neglected in the SST CCI uncertainty model. These include differences in the instantaneous field of view for channels of different wavelength, and local to regional variations in trace gas concentrations.

421 4. Validation of the Uncertainty Budget

Having constructed an initial uncertainty budget for remotely sensed 422 SSTs independently of in-situ data, we can now use these in-situ data to 423 validate our uncertainties (as well as the retrieved SST). In Section 3, we 424 characterised two quantifiable components of uncertainty relating to SSTs 425 calculated from satellite data at a pixel level (a random component due to 426 noise in the data and a locally systematic component arising from uncertain-427 ties varying on a synoptic scale within the retrieval) from which we construct 428 our initial uncertainty budget. We validate this budget using data from the 420 AATSR instrument spanning four years (2006 - 2009 inclusive) considering 430 both the N2 and D2 retrievals. The data used in the validation are taken 431 from the SST CCI multi-sensor match-up system (MMS) (Corlett et al., 432 2014) where drifting buoy and satellite observations are matched globally 433 under clear-sky conditions (Corlett et al., 2014). 434

Matches are filtered to include only the closest in-situ match in time to 435 the satellite observation and to check the quality of the in-situ data. Matches 436 can have a maximum time difference of 4 hours and maximum spatial sepa-437 ration of 10 km. Bad quality in-situ data are removed based on the following 438 criteria 1) absolute difference between NWP and in-situ SST greater than 439 10 K, 2) standard deviation of the in-situ SST history greater than 5 K and 440 3) standard deviation of the in-situ latitude history greater than 10 degrees. 441 Validation of satellite data using in-situ data necessitates a comparison be-442 tween a point measurement and the satellite footprint. There are uncertain-443 ties in this process arising from comparing two different types of observation 444 and geolocation errors in both the satellite and in-situ data. The filtering is 445

therefore necessary to minimise both spatial and temporal separation of the
satellite and in-situ observations (Minnett, 1986; Donlon et al., 2002; Corlett
et al., 2006).

For each match up, the uncertainties in the retrieved SST are calculated 449 as follows. The noise in a given observation is a function of both the channels 450 and associated brightness temperature, and is calculated by monitoring in-451 orbit blackbody temperature signals (Smith et al., 2012). For AATSR, the 452 NEdT is fairly consitent throughout the lifetime of the mission. These NEdT 453 values are are used to calculate the uncertainty due to uncorrelated effects 454 at L2 using the methodology presented in Section 3. The uncertainty from 455 locally systematic effects is quantified as a function of the TCWV consistent 456 with the banding of the retrieval coefficients. In both cases the uncertainties 457 are then propagated into the gridded product for validation of data in L3 458 format. For the gridded products, a sampling uncertainty is also calculated 459 due to the presence of cloud preventing observation of all pixels within a 460 given grid cell (Bulgin et al., 2016). This is an additional uncertainty due 461 to uncorrelated effects that is introduced in the gridding process. At both 462 the per pixel and gridded scales the uncertainty components are added in 463 quadrature to give a total uncertainty. 464

The validation data for the N2 and D2 pixel level retrievals are shown in the top two panels of Figure 4. Here we plot the standard deviation of the SST difference (retrieval minus drifting buoy) against the SST retrieval uncertainty which we have calculated independently represented by the thin black lines in Figure 4. The dashed lines indicate the uncertainty model we would expect to see based on retrieved SST minus drifting buoy differences.

There is a lower limit on this model of +/- 0.15 K which represents the 471 uncertainty in the drifting buoy measurements. We chose the time period 472 of 2006-2009 inclusive for our validation as the drifting buoy uncertainty 473 has been stable at around 0.15 K over this period (Lean and Saunders, 474 2012). The blue line on the plots indicate the median difference between the 475 retrieved and in-situ SST across all match-ups in each uncertainty bin (width 476 0.02 K). The standard error in this value is represented by the error bars. 477 Red lines at the end of the black bars indicate the statistical uncertainty in 478 the calculated standard deviation and are visible primarily for bins where 479 the number of contributing cases is small. 480

For the N2 pixel level data we find that our uncertainty estimates closely 481 match the expected uncertainty model below a total uncertainty of 0.25 K. 482 Above this threshold, our estimated retrieval uncertainties are too high: a 483 better fit would be obtained if the bins shifted to lower estimated uncertainty 484 values. For the D2 retrieval, we see that our uncertainties calculated within 485 the retrieval process show excellent agreement with the expected uncertainty 486 model. At a per-pixel level the dominant terms in the uncertainty budget 487 come from the uncorrelated and locally systematic effects, assuming that a 488 good cloud detection algorithm is used. Therefore the validation indicates 489 that our estimate of these components is well constrained. 490

We also consider the validation of uncertainties for gridded N2 and D2 retrievals across a 5x5 pixel domain approximately corresponding to 0.05°. In this case we also include the sampling uncertainty component in our initial uncertainty budget (Bulgin et al., 2016). The results for this validation are shown in the bottom two panels of Figure 4. When considering gridded data we find a larger range of estimated uncertainty than for the per pixel data. This is because SST varies across the gridded domain, and for cells that are not well sampled, the uncertainty on the mean SST increases. For the N2 gridded data we see a similar pattern to the N2 per pixel data with uncertainties being slightly overestimated. For the D2 gridded retrieval the overall uncertainties are smaller, but we underestimate the total uncertainty.

502 5. Discussion

Overall, we see that our independent uncertainty estimates show good 503 agreement with validation data using in-situ drifting buoy measurements. 504 The best agreement is for the D2 retrieval at a per-pixel level. For the N2 505 retrievals we see a similar over-estimation of uncertainties above 0.2-0.25506 K in both the pixel level and gridded products. The uncertainty budget 507 constructed is based on the errors that we currently have the capability to 508 estimate and propagate through the retrieval. Some of the sources of error 509 discussed in the earlier sections such as residual unscreened cloud contam-510 ination, failure to detect clear-sky pixels and aerosol are not yet included. 511 These may be larger across a gridded domain if they affect multiple pixels. 512

In this validation, the estimation of large scale systematic uncertainties has also been excluded, but in the SST CCI Version 1 products this is set to an invariant value of 0.1 K per pixel as a best estimate of the magnitude of this component, and then added in quadrature to the uncertainty budget (Merchant et al., 2014).

Although at present the uncertainty budget can not be fully constrained due to the limitations described in the Section 3, we are able to characterise

well the components resulting from random, locally systematic and sampling 520 effects across a range of retrievals for the ATSR instruments as evidenced 521 by the good validation statistics. On the relatively short spatial and tem-522 poral scales (pixel to gridded averages at 0.1° and instantaneous measure-523 ments) the uncertainties from uncorrelated and locall systematic effects are 524 the dominant terms in the uncertainty budget. The contributions from the 525 'missing' components are therefore relatively small under these SST retrieval 526 conditions. Empirical systematic effects (biases) are within the estimated un-527 certainties and these uncertainties can successfully distinguish more and less 528 certain SSTs. The approach outlined in this paper has a wider application to 529 coefficient based SST retrievals using other algorithms and data from other 530 instruments. If the data provider or user knows the NEdT distribution for 531 each channel used in the retrieval they can propagate this through the algo-532 rithm to obtain the uncertainty due to uncorrelated effects in the retrieved 533 SST. Data providers can use simulation studies to characterise the locally 534 systematic uncertainty in their retrieval scheme, and the sampling model is 535 applicable to any SST retrieval on the same spatial scales as discussed in 536 this paper provided that the uncertainty due to noise is removed first. Provi-537 sion of uncertainty information as part of the retrieval process then enables 538 validation of these uncertainty estimates, as well as the SST, using in-situ 539 data. 540

Figure 5 maps mean uncertainty estimates for 2010. The uncertainty maps show the square root of the mean of the error variance across all days with observations. Where more than one observation is available for a given day, the smallest error variance has been used. The uncertainty from uncorre-

lated effects (a) contains the noise and sampling uncertainty components and 545 when added to the uncertainty due to lcally systematic effects (b) in quadra-546 ture, produces the total uncertainty map (c). Total uncertainties typically 547 range between 0.1-0.25 K globally, with the highest values predominantly in 548 equatorial regions and some northern hemisphere high latitudes. The uncer-549 tainty due to uncorrelated effects is the larger contributor to this signal, and 550 in these regions scattered or patchy cloud cover increases sampling uncer-551 tainties. Figure 5 (d) also shows the ratio of the retrieved SST variability to 552 the uncertainty, calculated by dividing the standard deviation of the SST in 553 an given location over the whole of 2010 by the total uncertainty. The high-554 est ratios are seen in mid-latitude regions where SSTs show greater seasonal 555 variation. 556

557 6. Conclusions

In this paper we present a framework for the provision of uncertainty 558 estimates in coefficient based SST retrieval from satellite data, based on 550 propagation of noise, simulation of noise-free retrieval errors, and empirical 560 characterisation of sampling effects. The uncertainty estimates can be val-561 idated in their own right, in addition to validating the retrieved SST. We 562 provide a detailed discussion of different sources of uncertainty in the SST 563 retrieval and how to propagate these through the retrieval process. We derive 564 three uncertainty components here and in the companion paper; uncertain-565 ties due to uncorrelated, locally systematic and sampling effects. We apply 566 our derivation to AATSR data within the context of the SST CCI project 567 and find that our uncertainties validate well against in-situ data for both per 568

⁵⁶⁹ pixel and gridded products, and for two different retrieval algorithms.

570 7. Acknowledgements

The work undertaken in this paper was funded by the European Space Agency Sea Surface Temperature Climate Change Initiative project.

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Figure 1: Uncorrelated random errors and uncertainties in brightness temperature observations and SST retrieval. Panels a) and b) show simulated errors in the 11 and 12 μ m channels. Panels c) and d) show these errog6 propagated into SST retrievals for N2 and D2 retrievals. Panels e) and f) show the mean error at a 5x5 pixel resolution with a cloud mask superimposed on the data. Panels g) and h) show the associated uncertainty fields at a 5x5 pixel resolution.





37

Figure 2: AATSR retrieval errors for two different days from simulation studies (left and right). Plots show the difference between the 'true' and retrieved SST field. Plots in the upper panels show pressure contours hPa, and plots in the lower panels TCWV contours kg m⁻².



Figure 3: Uncertainties from a) locally systematic and b) uncorrelated effects as a function of total column water vapour for different channel combinations.



Figure 4: SST uncertainty validation against drifting buoy in-situ data. Top panels show pixel level uncertainties for N2 and D2 retrievals. Bottom panels show grid cell uncertainties (5x5 pixels approximately corresponding to a resolution of 0.05°) for N2 and D2 retrievals. Dashed lines show ideal uncertainty model accounting for uncertainties in the buoy data and geophysical uncertainties arising from a skin to depth comparison and colocation. Solid black lines show one standarged eviation of the retrieved minus buoy SST differences, and blue lines the median satellite minus buoy SST difference. Error bars show the standard error in these differences. Uncertainties in the retrieval uncertainty are indicated by red bars at the base and top of the solid black lines.



Figure 5: Annual means in SST retrieval uncertainties calculated from AATSR L3C data in 2010. Mean uncertainties are derived by adding all uncertainty observations in a given cell in quadrature, dividing by the number of observations and taking the square root. a) Shows uncertainty due to uncorrelated effects (noise and sampling uncertainty), b) shows noise due to locally systematic effects and c) total uncertainty. d) Shows the ratio of the SST standard deviation over 2010 to the total uncertainty.