

The characteristics of the magnetopause reconnection X-line deduced from lowaltitude satellite observations of cusp ions

Article

Supplemental Material

published correction to published version

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Correction to "The characteristics of the magnetopause reconnection X-line deduced from low-altitude satellite observations of cusp ions" by Lockwood et al.

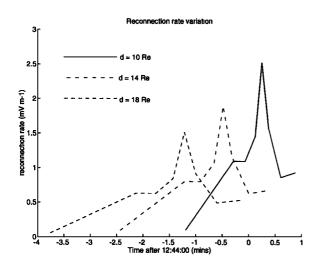
The version of the above paper printed in Vol. 21, No. 24 [Geophys. Res. Lett. 21, 2757-2760, 1994] was not the final one and was printed in error. There are two corrections.

(1). The derivation of figure 1 was incorrectly described; however, the figure itself is correct. Equation (1) and the second sentence below it should read:

$$d_n(m/2E)^{1/2} = t_s - (t_o + t_n) = (t_s V_s / V_c) - t_n$$
(1)

A field line opened at a time t_o is observed at a time t_s (in fig. 1 defined as zero at the OCB where $t_s = t_o$) at a distance $V_S t_s = V_C (t_s - t_o)$ poleward of the OCB.

(2). The correct version of figure 5 is given below and differs only in the labels for the axes. In the text, the average reconnection rate values quoted 5 lines from the bottom of column 1 on page 2760 should be 1.15, 0.85 and 0.67 mVm⁻¹ (for d_1 of 10, 14 and 18 R_E , respectively). Note that the corresponding mean and peak ionospheric flow speeds quoted in the subsequent 4 lines are correct.



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