

*Intercomparison of methods of coupling between convection and large-scale circulation. 2: comparison over non-uniform surface conditions*

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1 **Intercomparison of methods of coupling between**  
2 **convection and large-scale circulation. 2: Comparison**  
3 **over non-uniform surface conditions**

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4 **Key points.**

- 5     • Tropical convection  
6     • Large-scale parameterized dynamics

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7 **Abstract.** As part of an international intercomparison project, the weak  
8 temperature gradient (WTG) and damped gravity wave (DGW) methods  
9 are used to parameterize large-scale dynamics in a set of cloud-resolving mod-  
10 els (CRMs) and single column models (SCMs). The WTG or DGW method  
11 is implemented using a configuration that couples a model to a reference state  
12 defined with profiles obtained from the same model in radiative-convective  
13 equilibrium. We investigated the sensitivity of each model to changes in SST,  
14 given a fixed reference state. We performed a systematic comparison of the  
15 WTG and DGW methods in different models, and a systematic comparison  
16 of the behavior of those models using the WTG method and the DGW method.

17 The sensitivity to the SST depends on both the large-scale parameteriza-  
18 tion method and the choice of the cloud model. In general, SCMs display a  
19 wider range of behaviors than CRMs. All CRMs using either the WTG or  
20 DGW method show an increase of precipitation with SST, while SCMs show  
21 sensitivities which are not always monotonic. CRMs using either the WTG  
22 or DGW method show a similar relationship between mean precipitation rate  
23 and column-relative humidity, while SCMs exhibit a much wider range of be-  
24 haviors. DGW simulations produce large-scale velocity profiles which are smoother  
25 and less top-heavy compared to those produced by the WTG simulations.

26 These large-scale parameterization methods provide a useful tool to iden-  
27 tify the impact of parameterization differences on model behavior in the pres-  
28 ence of two-way feedback between convection and the large-scale circulation.

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## 1. Introduction

29 A key issue in understanding the tropical climate and its variability is the understand-  
30 ing of the two-way interaction between tropical deep convection and large-scale tropical  
31 circulations. Numerical models which simultaneously simulate convection and large-scale  
32 circulations are computationally expensive due to the large range of spatial scales between  
33 individual convective cells and large-scale tropical circulations. Some examples include  
34 large-domain, high-resolution simulations as those conducted in projects such as Cascade  
35 [e.g., *Holloway et al.*, 2012] and the global cloud-resolving modeling using Nonhydrostatic  
36 ICosahedral Atmosphere Model [e.g., *Miura et al.*, 2005].

37 Many single column model (SCM) and cloud-resolving model (CRM) studies have simu-  
38 lated the interactions of tropical deep convection with a prescribed large-scale flow, possi-  
39 bly based on idealization or experimental campaign [e.g., *Tompkins*, 2001; *Xu et al.*, 2002;  
40 *Derbyshire et al.*, 2004; *Petch et al.*, 2006]. In such studies, the time scale characterizing  
41 changes in convection is assumed to be short compared to the time scale characterizing  
42 changes in the large-scale flow. Simulations with predefined large-scale flow have provided  
43 much useful insight. However, the precipitation rates produced are too much constrained  
44 due to the predefined large-scale moisture advection [*Mapes*, 1997; *Sobel and Bretherton*,  
45 2000] and thus, such simulations cannot be used to understand the factors that con-  
46 trol the occurrence and intensity of tropical deep convection [*Sobel et al.*, 2004]. On the  
47 other hand, in non-equilibrium conditions there is a close link between convection and

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48 the large-scale flow such that ignoring the feedback of convection on the large-scale flow  
49 is not appropriate [*Mapes, 1997; Holloway and Neelin, 2010; Masunaga, 2012*].

50 The two-way interaction between tropical deep convection and large-scale tropical flow  
51 has been studied at a reasonable computational cost in both SCMs and CRMs using  
52 various forms of parameterized large-scale dynamics. This study compares two methods  
53 of parameterized large-scale dynamics—the weak-temperature gradient (WTG) method  
54 and the damped gravity wave (DGW) method—in a set of CRMs and SCMs.

55 The WTG method derives the large-scale vertical velocity from buoyancy anomalies. It  
56 has been applied to parameterize large-scale tropical circulations that either consume the  
57 simulated heating and accordingly maintain zero horizontal temperature gradient [*Sobel*  
58 *and Bretherton, 2000*] or remove the horizontal temperature gradient over a short but  
59 nonzero time-scale [e.g., *Raymond and Zeng, 2005; Sessions et al., 2010; Daleu et al.,*  
60 *2012; Sessions et al., 2015*]. A recent innovation of the WTG method involves spectral  
61 decomposition of heating in the vertical dimension [*Herman and Raymond, 2014*]. The  
62 DGW method derives the large-scale vertical velocity directly from the approximated mo-  
63 mentum equations. It has been applied to study the two-way coupling between convection  
64 and large-scale dynamics, with the latter being simplified to a linear gravity wave of a  
65 single horizontal wavenumber [*Kuang, 2008, 2011; Wang et al., 2013; Romps, 2012a, b;*  
66 *Edman and Romps, 2015*].

67 In the simulations using the WTG or the DGW method the large-scale forcing diagnosed  
68 from the domain-mean temperature anomalies induces a moisture source. Therefore, tra-  
69 ditional intercomparisons with prescribed large-scale forcing (e.g., TOGA COARE and

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70 DYNAMO) and intercomparisons in which moisture source is a relaxation to a prescribed  
71 profile [*Derbyshire et al.*, 2004] are extended here to simulations in which convection within  
72 the simulated domain feeds back on the large-scale forcing which in turns drives moisture  
73 advection. The implementation of the WTG and DGW methods has always used a con-  
74 figuration that couples a simulated column to a reference state [e.g., *Raymond and Zeng*,  
75 2005; *Sobel et al.*, 2007; *Sessions et al.*, 2010; *Wang and Sobel*, 2011; *Kuang*, 2008, 2011;  
76 *Wang and Sobel*, 2012; *Wang et al.*, 2013; *Romps*, 2012a, b] until recently when *Daleu*  
77 *et al.* [2012] developed a new configuration that couples two simulated columns via a  
78 WTG-derived large-scale circulation [*Daleu et al.*, 2012, 2014]. Much insight has been  
79 learned from these efforts. Unfortunately, many aspects of the large-scale parameteriza-  
80 tion methods remain uncertain since results using these two large-scale parameterization  
81 methods show both similarities and discrepancies in model behavior.

82 In order to understand the different behaviors of these large-scale parameterization  
83 methods, this international intercomparison project—the GASS-WTG project—was devel-  
84 oped by the Global Energy and Water Exchanges (GEWEX) Global Atmospheric Systems  
85 Modelling Panel (GASS). The goals of this project are to develop community understand-  
86 ing of the WTG and DGW methods, to identify differences in behavior of SCMs compared  
87 to CRMs to inform parameterization development, and to assess the usefulness of these  
88 approaches as tools for parameterization development. In this study, we will evaluate  
89 the CRMs and SCMs by comparing the strengths of the diagnosed large-scale forcing  
90 and the precipitation rates which result from both the model physics and the parame-  
91 terized large-scale dynamical feedback. These two-way feedbacks between convection and

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92 the large-scale forcing will helps us to identify weaknesses in our SCM parameterization  
93 schemes and their likely behaviors in general circulation models. However, such compar-  
94 ison will be helpful only if a greater consistency is obtained among CRMs than among  
95 SCMs.

96 In Part 1 of this study [*Daleu et al.*, 2015], the aim was to understand what causes  
97 discrepancies in model behavior when surface conditions in the simulated column are  
98 identical to those of the reference state. We implemented the WTG and DGW methods  
99 in a set of CRMs and SCMs. For each model, the reference state was defined from  
100 profiles obtained in the radiative-convective equilibrium (RCE) simulation of that model.  
101 WTG and DGW simulations were performed with the same SST as in the reference state  
102 and were initialized with profiles from the reference state. Some models produced an  
103 equilibrium state which was almost identical to the corresponding RCE reference state.  
104 In contrast, other models developed a large-scale circulation which resulted in either  
105 substantially higher or lower precipitation rates in the simulated column compared to  
106 the implied value for the RCE reference column. We also explored the sensitivity of  
107 the final equilibrium state to the initial moisture conditions. We found that while some  
108 models are not sensitive to the initial moisture conditions (independent of the method  
109 used to parameterize the large-scale circulation), other models may support two distinct  
110 precipitating equilibrium states using either the DGW or WTG method. We also found  
111 that some models using the WTG method (but not using the DGW method) can support

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112 either an equilibrium state with persistent, precipitating convection or an equilibrium  
113 state with zero precipitation.

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114 *Daleu et al.* [2015] revealed some weaknesses of the WTG method. For instance, over  
115 uniform SST, the existence of the nonprecipitating equilibrium state in some models  
116 was sensitive to the choice of the parameters used in the WTG calculations (e.g., the  
117 nominal boundary layer depth). In addition, DGW simulations over uniform SST and  
118 with nearly uniform radiative forcing were more likely to reproduce the RCE reference  
119 conditions and produced large-scale pressure velocities which were smoother compared  
120 to those produced by the WTG simulations. Aside from the choice of the large-scale  
121 parameterization method and the details of its implementation, various other factors in  
122 the convective models were important for the evolution of convection and its interactions  
123 with parameterized large-scale dynamics. For instance, we found that CRMs using either  
124 the WTG or DGW method produced broadly similar results, while SCMs produced a  
125 much wider range of behaviors.

126 Whilst *Daleu et al.* [2015] considered the case where the simulated column had the  
127 same SST as the RCE reference state, this paper focuses on the sensitivity to the SST  
128 in the simulated column, which has been a major focus of previous studies using these  
129 approaches [e.g., *Raymond and Zeng, 2005; Sobel et al., 2007; Wang and Sobel, 2011*].  
130 *Daleu et al.* [2015] used the term “Uniform SST” to refer to conditions in which the  
131 simulated column has the same SST as in the RCE reference state. In the present study,  
132 we use the same set of CRMs and SCMs presented in *Daleu et al.* [2015] and we use  
133 the term “Non-uniform SST” to refer to conditions in which the simulated column has a  
134 value of SST which is different to that of the RCE reference state. For each model, we  
135 fix the reference state and perform a series of WTG and DGW simulations with a range  
136 of SSTs in the simulated column. We perform a systematic comparison of the WTG and

137 DGW methods with a consistent implementation in the models, and also a systematic  
138 comparison of the behavior of the models given the same large-scale parameterization  
139 method.

140 This paper is organized as follows. Section 2 briefly describes the models that have  
141 contributed to this study. Section 3.1 outlines our implementation of the WTG and DGW  
142 methods (full details are available in *Daleu et al.* [2015]), while Section 3.2 describes the  
143 configurations of our numerical simulations. Section 4 compares the results of the WTG  
144 and DGW simulations over non-uniform SSTs. Finally, the conclusions and implications  
145 of our study are discussed in section 5.

## 2. Description of models

146 Six groups participating in this intercomparison study performed simulations with the  
147 same set of CRMs and SCMs presented in *Daleu et al.* [2015]. The models are listed in  
148 Tables 1 and 2 for CRMs and SCMs, respectively.

### 2.1. Cloud Resolving Models

149 There are five CRMs, including two in three-dimensions [3-D] and three in two-  
150 dimensions [2-D]. The 3-D CRMs are the Weather Research and Forecast model version  
151 3.3 (WRF) [*Skamarock et al.*, 2008] and the mesoscale, nonhydrostatic atmospheric model  
152 (MesoNH) [*Lafore et al.*, 1997]. The 2-D CRMs are the Langley Research Center Cloud-  
153 Resolving Model (LaRC-CRM) [*Cheng and Xu*, 2006], the New Mexico Tech cloud model  
154 version 3 (NMTCMv3) introduced in *Raymond and Zeng* [2005], with modifications and  
155 enhancements described in *Herman and Raymond* [2014], and the Met Office Large Eddy

156 Model at version 2.4 (LEMv2.4) [*Shutts and Gray, 1994; Petch and Gray, 2001*]. The  
157 reader is referred to *Daleu et al. [2015]* for a more complete description of these CRMs.

## 2.2. Single-Column Models

158 Two pairs of the SCMs come from different versions of the same model. One of the pairs,  
159 LMDzA and LMDzB, are the SCM versions of the atmospheric components of IPSL-CM5A  
160 and IPSL-CM5B [*Dufresne et al., 2013*]. The other pair, EC-Earthv1 and EC-Earthv3,  
161 are SCMs based on the atmospheric general circulation model IFS, cycles 31r1 and 36r4  
162 respectively of the European Centre for Medium-Range Weather Forecasts (ECMWF)  
163 [*Hazeleger et al., 2010*]. ARPV6 is the SCM version of the atmospheric component of the  
164 CNRM-CM, an updated version from that used in CMIP5 [*Voldoire et al., 2013*], GISS-  
165 SCM is the SCM version of the National Aeronautics and Space Administration Goddard  
166 Institute for Space Studies, an updated version from that used in CMIP5 [*Schmidt et al.,*  
167 2014], and UMv7.8 is the SCM version of the UK Met Office Unified Model [*Davies et al.,*  
168 2005]. The reader is referred to *Daleu et al. [2015]* for a more complete description of  
169 these SCMs.

## 2.3. Overall approach

170 The CRMs have horizontal domain sizes ranging between 128 and 256 km and hori-  
171 zontal resolution ranging between 0.5 and 4 km. The lateral boundary conditions are  
172 periodic for all prognostic variables in all CRMs. For CRMs in 2-D, the domain-mean  
173 wind speeds in the along-domain direction and in the across-domain direction are relaxed  
174 toward vertically uniform values of 0 and 5 m s<sup>-1</sup>, respectively, both with a relaxation  
175 time-scale of 6 h. For fair comparison of 2-D CRM simulations with 3-D CRM simulations

and with SCM simulations, the horizontal domain-mean wind speed components in the 3-D CRMs and SCMs are also relaxed toward vertically uniform values of 0 and 5 m s<sup>-1</sup>.

For all of these models, the lower boundary condition is a spatially uniform and time-independent SST, and the Coriolis force is zero. We force each model with the idealized cooling profile defined in *Daleu et al.* [2015]. The tendency of temperature due to radiative cooling,  $(\partial T/\partial t)_{RC}$ , is homogeneous and non-interactive throughout most of the troposphere, and it acts to maintain the temperature toward a fixed value of 200 K at levels with  $\bar{p} < 100$  hPa, with a relaxation time scale  $\alpha_T^{-1} = 1$  day. That is,

$$\left(\frac{\partial T}{\partial t}\right)_{RC} = \begin{cases} -1.5 & \text{if } \bar{p} \geq 200 \\ -1.5 \left(\frac{\bar{p}-100}{100}\right) - \alpha_T \left(\frac{200-\bar{p}}{100}\right) (\bar{T} - 200) & \text{if } 100 < \bar{p} < 200. \\ -\alpha_T (\bar{T} - 200) & \text{if } \bar{p} \leq 100 \end{cases} \quad (1)$$

### 3. Parameterization of the large-scale dynamics and experiment setup

#### 3.1. Parameterization of the large-scale dynamics

In the present study, the large-scale circulation is parameterized using two methods: the WTG and DGW methods. As in *Daleu et al.* [2015], the implementation of the WTG or DGW method involves an interactive column that is coupled to a reference state.

A full description of the implementation of the WTG method is given in *Daleu et al.* [2015]. The large-scale pressure velocity,  $\bar{\omega}$  between 850 and 100 hPa acts to reduce the difference in the domain-mean virtual potential temperature between the simulated column and the reference state,  $\bar{\theta}_v - \bar{\theta}_v^{Ref}$ , over a specified time-scale,  $\tau$ . That is,

$$\bar{\omega} \frac{\partial \bar{\theta}_v^{Ref}}{\partial p} = \frac{\bar{\theta}_v - \bar{\theta}_v^{Ref}}{\tau}. \quad (2)$$

Above 100 hPa  $\bar{\omega}$  is set to zero. Below the nominal boundary layer top, 850 hPa, we calculate the values of  $\bar{\omega}$  by linear interpolation in pressure from the value diagnosed at the

195 first model level above 850 hPa to zero at the surface. Experiments to assess sensitivities  
 196 of the final equilibrium state to the depth of the boundary layer are presented in *Daleu*  
 197 *et al.* [2015].

198 A full description of the implementation of the DGW method is given in *Daleu et al.*  
 199 [2015]. The second-order derivative of  $\bar{\omega}$  is related to the difference in the domain-mean  
 200 virtual temperature between the simulated column and the reference state,  $\bar{T}_v - \bar{T}_v^{Ref}$ , as

$$201 \quad \frac{\partial}{\partial p} \left( \epsilon \frac{\partial \bar{\omega}}{\partial p} \right) = \frac{k^2 R_d}{\bar{p}^{Ref}} (\bar{T}_v - \bar{T}_v^{Ref}), \quad (3)$$

202 where  $R_d$  is the gas constant of dry air.  $\epsilon$  and  $k$  are the mechanical damping coefficient  
 203 and the horizontal wavenumber, respectively.

204 As in *Daleu et al.* [2015], the large-scale circulation parameterized using either equation 2  
 205 or 3 introduces additional source and sink terms to the potential temperature and water  
 206 vapor equations only. The prognostic equation for potential temperature includes the  
 207 tendency due to vertical advection by the parameterized large-scale circulation. That is,

$$208 \quad \left( \frac{\partial \theta}{\partial t} \right)_{LS} = -\bar{\omega} \frac{\partial \bar{\theta}}{\partial p}. \quad (4)$$

209 The prognostic equation for specific humidity of water vapor ( $q_v$ ) also includes the large-  
 210 scale tendency due to vertical advection, as well as an additional contribution representing  
 211 the horizontal advection of the reference state air into the simulated domain by the pa-  
 212 rameterized large-scale circulation. That is,

$$213 \quad \left( \frac{\partial q_v}{\partial t} \right)_{LS} = -\bar{\omega} \frac{\partial \bar{q}_v}{\partial p} + \max \left( \frac{\partial \bar{\omega}}{\partial p}, 0 \right) (\bar{q}_v^{Ref} - \bar{q}_v). \quad (5)$$

### 3.2. Experiment Setup

214 For each model, a radiative-convective equilibrium (RCE) simulation (no large-scale  
 215 parameterized dynamics) is first performed over an SST of 300 K. The mean thermody-

216 namic profiles at equilibrium in that simulation are used to define the reference state of  
217 that model. We keep the reference state fixed and investigate the sensitivity of the final  
218 equilibrium state to the SST in the simulated column as in *Wang and Sobel* [2011].

219 For each of the models listed in Tables 1 and 2, we performed the WTG and DGW  
220 simulations of a colder column (using SSTs of 298 and 299.5 K), a warmer column (using  
221 SSTs of 300.5, 301, 301.5 and 302 K), and over a uniform SST (using an SST of 300 K;  
222 results presented in *Daleu et al.* [2015]). The adjustment time-scale used in the WTG  
223 calculations is  $\tau = 3$  h. In the DGW calculations, we fix the value of  $\epsilon$  to  $1 \text{ day}^{-1}$  and  
224 solve equation 3 with a single horizontal wavenumber  $k = 10^{-6} \text{ m}^{-1}$ . These are typical  
225 values used in previous WTG and DGW studies [e.g., *Herman and Raymond*, 2014; *Daleu*  
226 *et al.*, 2012; *Wang and Sobel*, 2011; *Wang et al.*, 2013], including *Daleu et al.* [2015]. They  
227 have been chosen such that the WTG simulation and the corresponding DGW simulation  
228 produce large-scale circulations that are comparable in strength for similar temperature  
229 anomalies. The calculations of  $\bar{w}$  given by equations 2 and 3 are performed either every  
230 10 min (for models with integration time steps smaller or equal to 10 min) or at every  
231 model time step (for models with integration time steps greater than 10 min).

232 The results presented in *Daleu et al.* [2015], and in other previous studies [e.g., *Sobel*  
233 *et al.*, 2007; *Sessions et al.*, 2010] show that some SCMs and CRMs using the WTG method  
234 can sustain either a dry equilibrium state or a precipitating equilibrium state, given suffi-  
235 ciently different initial moisture conditions (known as multiple equilibria). Therefore, it is  
236 possible that some of our WTG simulations that exhibit precipitating equilibrium states  
237 would instead result in dry equilibrium states if initialized with very dry moisture condi-  
238 tions. Multiple equilibria and their dependence on parameters in the WTG calculations

239 have already been investigated in *Daleu et al.* [2015], and they are outside the scope of  
240 the present paper.

241 The WTG and DGW calculations are initialized with profiles from the models' RCE  
242 reference state at 300 K and are allowed to evolve until a new quasi-equilibrium state with  
243 parameterized large-scale circulation is reached. The RCE reference profiles differ from  
244 model to model, with large differences obtained among SCMs (see Figure 3 in *Daleu et al.*  
245 [2015]). The value of surface sensible heat flux also differs between models (not shown)  
246 but is much smaller than surface latent heat flux, such that the main balance in the RCE  
247 state is between the precipitation rate and the column-integrated radiative cooling rate.  
248 Due to the dependence of radiative cooling profile on temperature above 200 hPa (see  
249 equation 1), the value of column-integrated radiative cooling rate differs from model to  
250 model. The values of mean precipitation rate obtained in the RCE simulations with an  
251 SST of 300 K are summarized in the last rows of Tables 1 and 2 for CRMs and SCMs,  
252 respectively.

253 We conducted a set of WTG and DGW simulations over non-uniform SSTs using each  
254 of the models listed in Tables 1 and 2. The simulations are integrated over different  
255 periods of time ranging between 50 and 250 days, as the time-scale of adjustment to a  
256 quasi-equilibrium state with the parameterized large-scale circulation differs from model to  
257 model and also depends on which large-scale parameterization method is used. The quasi-  
258 equilibrium state is reached when a statistically steady state temperature and humidity  
259 profiles are achieved when averaged over a long period of time. The mean states and  
260 statistics at equilibrium of the simulations to be discussed have been obtained by averaging



261 over the last 20 days in 50-day simulations, 30 days in 100-day simulations, and 100 days  
 262 in 250-day simulations.

## 4. Results

263 In this section, we present the profiles of large-scale pressure velocity and the mean  
 264 precipitation rates at equilibrium for different values of SST in the simulated column.  
 265 We also present the mean precipitation rates, circulation strength, and column-relative  
 266 humidity in a set of scatter plots.

### 4.1. Parameterized large-scale circulation and mean precipitation rates

267 Figures 1 and 2 show the profiles of  $\bar{w}$  obtained at equilibrium in the WTG and DGW  
 268 simulations, respectively. Results are shown for all models listed in Tables 1 and 2 and  
 269 for SSTs of 298, 299.5, 300 K (uniform SST; results presented in *Daleu et al.* [2015]),  
 270 300.5, 301, 301.5 and 302 K. For models in height coordinates, we expressed the large-  
 271 scale vertical velocities in  $\text{Pa s}^{-1}$  by applying the factor “ $-\rho g$ ,” where  $\rho$  is density and  $g$   
 272 is the gravitational acceleration.

273 To provide a more quantitative evaluation of the WTG and DGW simulations, we  
 274 calculated the ratio of mean precipitation rate in the simulated column,  $P$ , to the value  
 275 of the corresponding RCE reference state,  $P_{Ref}$ . We also calculated the mass-weighted  
 276 vertical integral of the large-scale pressure velocities presented in Figures 1 and 2;  $\Omega =$   
 277  $\int \bar{w} dp / \Delta p$ , where  $\Delta p$  is the depth of the troposphere. The numerical values of  $\Omega$  and  
 278  $P/P_{Ref}$  are listed in Tables 3 and 4 for CRMs and SCMs, respectively. Figure 3 shows  
 279  $P/P_{Ref}$  as a function of the SST in the simulated column, and Figure 4 shows scatter  
 280 plots of  $\Omega$  versus  $P/P_{Ref}$  for all SSTs.

281 **4.1.1. Variations between models**

282 For a given SST in the simulated column, the characteristic vertical structure of the  
283 large-scale circulation at equilibrium differs from model to model, and it also depends  
284 on the large-scale parameterization method used. Over an SST of 302 K (red curves in  
285 Figures 1 and 2), for example, models using the WTG method exhibit a range of large-  
286 scale pressure velocity profiles which vary from unimodal ascent through the column with  
287 very top-heavy profiles (e.g., WRF; Figure 1a), to more uniform unimodal profiles (e.g.,  
288 LaRC-CRM; Figure 1c), to bi-modal profiles (e.g., EC-Earthv1; Figure 1k), to profiles  
289 with distinct minima near the freezing level (e.g., UMv7.8; Figure 1j), including some  
290 with weak descent near the freezing level (e.g., GISS-SCM; Figure 1h). As seen in *Daleu*  
291 *et al.* [2015], the DGW method produces large-scale pressure velocity profiles which are  
292 smoother than those produced using the WTG method (compare Figures 1 and 2).

293 Over cold SSTs (298 and 299.5 K), some models produce large-scale pressure velocity  
294 profiles which are insensitive to the SST. In such simulations, convection is inhibited com-  
295 pletely and the heating due to the diagnosed large-scale circulation balances the prescribed  
296 radiative cooling. Some examples are the WTG simulations of LEMv2.4 with SSTs of 298  
297 and 299.5 K which produce zero precipitation rates (see Table 3) and indistinguishable  
298 large-scale pressure velocity profiles (see dark blue and light blue curves in Figure 1e).

299 Over warm SSTs, the large-scale pressure velocity profiles and precipitation rates are  
300 sensitive to the SST in all the models using either the WTG or DGW method. The  
301 sensitivity differs from model to model, and there is much diversity even among CRMs.  
302 Using the DGW method, for example, the two 3D CRMs (WRF and MesoNH) with an  
303 SST of 302 K produced large-scale pressure velocities and precipitation rates which differ

304 by more than a factor of two (compare the red curves in Figures 2(a) and 2(b), and the  
305 values of  $P/P_{Ref}$  in Table 3). However, all CRMs with an SST  $\geq 301$  K have large-scale  
306 pressure velocities increasing upward to around 400 hPa using the DGW method and to  
307 around 250 hPa using the WTG method.

308 The large-scale pressure velocity profiles produced in most SCM simulations vary con-  
309 siderably from the very top-heavy profiles (e.g., GISS-SCM using the DGW method, see  
310 Figure 2h) through weakly top-heavy profiles (e.g., LMDzB using the WTG method, see  
311 Figure 1g) to the bottom-heavy profiles (e.g., EC-Earthv1 using the WTG method; see  
312 Figure 1k), and some of the pressure velocity profiles show very detailed structures in the  
313 vertical (e.g., UMv7.8 using the WTG method; see Figure 1j). Similar to the results of  
314 *Wang et al.* [2013], the pressure velocity profiles produced using the DGW method are  
315 much smoother and tend to be slightly less top-heavy compared to those produced using  
316 the WTG method (compare Figures 1 and 2).

#### 317 4.1.2. Variations with SST

318 The impact of the SST is readily seen. At SST= 298 K, all the models using either the  
319 WTG or DGW method produce uniform large-scale descent (see the dark blue curves in  
320 Figures 1 and 2). In some of these simulations, the large-scale circulation inhibits pre-  
321 cipitating convection completely (e.g., NMTCMv3 using the DGW method; see Table 3),  
322 while in others an equilibrium state with light precipitation can be achieved (e.g., LMDzB  
323 using the WTG method; see Table 4).

324 At SST= 299.5 K, all CRMs using either the WTG or DGW method produced uniform  
325 large-scale descent. With the exception of GISS-SCM using the WTG method, which  
326 produces large-scale ascent in the upper troposphere (light blue curve in Figure 1h), the

327 SCMs produce either a uniform large-scale descent throughout the column (e.g., ARPV6  
 328 using the WTG method; light blue curve in Figure 1i) or large-scale descent in the upper  
 329 troposphere and a very weak circulation in the lower troposphere (e.g., EC-Earthv3 using  
 330 the WTG method; light blue curve in Figure 1l). The WTG and DGW simulations which  
 331 produce uniform large-scale descent result in very low precipitation compared to the value  
 332 of the RCE reference state, consistent with the negative moisture transport implied by  
 333 the resulting large-scale circulation (e.g., MesoNH using the WTG method; see Table 3),  
 334 with some simulations producing zero precipitation at equilibrium (e.g., WRF using the  
 335 WTG method; see Table 3). The WTG and DGW simulations which produce large-scale  
 336 descent in the upper troposphere and a very weak circulation in the lower troposphere are  
 337 dominated by shallow convection and thus, result in smaller reductions in precipitation  
 338 compared to the value of the RCE reference state (e.g., EC-Earthv3 using the WTG  
 339 method; see Table 4). However, in the WTG simulation of GISS-SCM with an SST of  
 340 299.5 K the mean precipitation rate at equilibrium is slightly increased (with respect to  
 341 the value of the RCE reference state) to balance the net small cooling produced by the  
 342 large-scale ascent in the upper troposphere. In contrast, the DGW simulation of GISS-  
 343 SCM with an SST of 299.5 K produces a different sign of the circulation with a reduction  
 344 of precipitation (see Table 4).

345 The results of the WTG and DGW simulations over uniform SST are presented in  
 346 *Daleu et al.* [2015]. There, we considered that a WTG or DGW simulation over a uniform  
 347 SST replicated the corresponding RCE reference state to a good approximation if  $0.9 <$   
 348  $P/P_{Ref} < 1.1$  and  $-0.4 \times 10^{-2} < \Omega < 0.4 \times 10^{-2} \text{ Pa s}^{-1}$ . The values of  $\Omega$  and  $P/P_{Ref}$   
 349 for such simulations are both bold-faced in Tables 3 and 4. Some models replicate the

350 corresponding RCE reference state to a good approximation. In contrast, other models  
351 sustain a large-scale ascent (or descent) which results in substantially higher (or lower)  
352 precipitation rate in the simulated column compared to the value of the corresponding  
353 RCE reference state.

354 Those models which produce a lower precipitation rate over a uniform SST of 300 will  
355 not produce a mean precipitation rate which is equivalent to the value of the RCE reference  
356 state unless the SST in the simulated column is increased, consistent with the results of  
357 *Raymond and Zeng* [2005]. An example is UMv7.8 using the WTG method (see  $P/P_{Ref}$   
358 as a function of the SST; green curve in Figure 3b). Similarly, models which produce a  
359 higher precipitation rate will not produce a mean precipitation rate which is equivalent to  
360 the value of the RCE reference state unless the SST in the simulated column is decreased  
361 (e.g., AR Pv6 using the WTG method; solid black curve in Figure 3b).

362 An SST of 300.5 K results in substantially higher precipitation rate ( $P/P_{Ref} > 1.1$ )  
363 in all the WTG and DGW simulations, except EC-Earthv1. A large proportion of these  
364 simulations produce uniform large-scale ascent (e.g., GISS-SCM using the DGW method,  
365 dark green curve in Figure 2h). Other simulations produce large-scale circulations with  
366 a layer of descent near the freezing layer, but which nonetheless result in net column-  
367 integrated cooling and moistening of the simulated column (e.g., AR Pv6 using the WTG  
368 method, dark green curve in Figure 1i and  $P/P_{Ref} > 1.1$  in Table 4). In contrast, the  
369 WTG and DGW simulations of EC-Earthv1 with an SST of 300.5 K produce large-scale  
370 circulations with ascent in the upper troposphere and descent in the lower troposphere  
371 (dark green curves in Figures 1k and 2k), despite producing ascent in the lower troposphere  
372 over a uniform SST of 300 K (black curves in Figures 1k and 2k). In this model using

373 the DGW method, the large-scale circulation cools and moistens the upper troposphere  
374 at the same rates as it warms and dries the the lower troposphere. As a result, the  
375 column-integrated heating and moistening rates produced by the large-scale circulation  
376 are both negligible and thus, the simulated column achieves an equilibrium precipitation  
377 rate which is very close to the corresponding RCE reference state (see value of  $P/P_{Ref}$   
378 in Table 4). In contrast, using the WTG method the upper tropospheric cooling and  
379 moistening do not prevent a reduction in precipitation rate due to the lower tropospheric  
380 warming and drying (see Table 4). A similar result is obtained in the WTG simulation  
381 of EC-Earthv1 with an SST of 301 K (see the light green curve in Figure 1k and and the  
382 value of  $P/P_{Ref}$  in Table 4). The WTG and DGW simulations of EC-Earthv1 with an  
383 SST of 301 K produce different signs of the integrated circulation.

384 At SSTs  $> 301$  K, the mean precipitation rate is increased compared to the value  
385 of the corresponding RCE reference state in all the models using either the WTG or  
386 DGW method. These simulations produce uniform large-scale ascent in the simulated  
387 column, with the exceptions of the WTG simulations of ARPV6 and GISS-SCM, in which  
388 a thin layer of descent between 750 and 650 hPa does not prevent an increase in mean  
389 precipitation rate.

390 For all CRMs using either the WTG or the DGW method the simulated column evolves  
391 toward a new quasi-equilibrium state with mean precipitation rate increasing non-linearly  
392 with SST, consistent with SCM results from *Sobel and Bretherton* [2000], and *Ramsay*  
393 *and Sobel* [2011]. In contrast, the SCMs show sensitivities of the mean precipitation rate  
394 to the SST which are not always monotonic (e.g., EC-Earthv1 using either the WTG or  
395 DGW method; solid red curves in Figures 3b and 3d).

396 Within an individual model, the sensitivity of precipitation rate to the SST depends  
 397 on which large-scale parameterization method is used. An example is WRF which shows  
 398 a stronger sensitivity under the DGW method than under the WTG method (compared  
 399 the dashed curves in Figures 3a and 3c). On the other hand, given one of the large-  
 400 scale parameterization methods (either the WTG or DGW method), the sensitivity of  
 401 precipitation rate to the SST differs from model to model.

402 An approximately linear relationship between  $\Omega$  and the mean precipitation rate is  
 403 expected, since the mean vertical motion and mean vertical moisture advection are cor-  
 404 related. In our study, despite the differences in the pressure velocity profiles,  $\Omega$  and the  
 405 mean precipitation rate show a fairly linear relationship (see Figure 4) and only models  
 406 with unusual vertical pressure velocity profiles shows deviations from this linear relation-  
 407 ship (e.g., GISS-SCM using the WTG method; see Figure 1h and circles in Figure 4b).  
 408 Most of the models meet the expectation that the large-scale circulation and precipitation  
 409 rate should increase with SST. Models which show a monotonic increase of precipitation  
 410 with SST also show a monotonic increase of precipitation with  $\Omega$  (WRF using the WTG  
 411 method; dashed curve with solid circles in Figure 3a and solid circles in Figure 4a). In con-  
 412 trast, models which show a non monotonic increase of precipitation with SST also show a  
 413 non monotonic increase of precipitation with  $\Omega$  (e.g., GISS-SCM using the WTG method  
 414 at warm SSTs; dashed black curve with circles in Figure 3b and circles in Figure 4b).

## 4.2. Precipitation and Column relative humidity

415 In this section, we examine the relationship between precipitation and the column rela-  
 416 tive humidity (hereafter  $CRH$ ) in our WTG and DGW simulations.  $CRH$  is calculated as  
 417 the ratio of column-integrated water vapor to its saturation value. Figure 5 shows scatter

418 plots of  $P$  versus  $CRH$ . It also shows the exponential fit for the observed monthly mean  
 419 precipitation over the tropical oceans obtained by *Bretherton et al.* [2004](solid curve).

420 That is

$$421 \quad P(\text{mmd}^{-1}) = \exp[11.4(CRH - 0.522)]. \quad (6)$$

422 To account for the variations in  $CRH$  of the RCE reference state, we also consider Figure 6,  
 423 which shows scatter plots of the ratios  $P/P_{Ref}$  versus  $CRH/CRH_{Ref}$ , where  $CRH_{Ref}$  is  
 424 the column-integrated relative humidity of the RCE reference state. The values of  $P$  and  
 425  $CRH$  are those obtained at equilibrium in the WTG and DGW simulations of each of the  
 426 models listed in Tables 1 and 2 with the values of SST ranging between 298 and 302 K.

427 Generally, the mean precipitation rate increases as  $CRH$  increases, except in the DGW  
 428 simulations of LMDzA with SSTs  $\leq 300.5$  K in which  $CRH$  decreases while precipitation  
 429 rate increases (see left facing triangles in Figure 5d). The decrease of  $CRH$  with mean  
 430 precipitation rate is unusual, but we do not investigate this further in this study.

431 In a large proportion of the models, there is a threshold value of  $CRH$  below which  
 432 there is virtually no precipitation or strongly reduced precipitation rate (with respect to  
 433 the value of the RCE reference state) and above which precipitation rate rapidly increases  
 434 with  $CRH$ . Below this threshold, the WTG and DGW simulations show changes in  
 435 mean precipitation rate that are relatively small for large changes in  $CRH$ . Above this  
 436 threshold, a significant increase in precipitation rate is obtained, followed by a sharp  
 437 pickup of mean precipitation rate as  $CRH$  increases further. The value of this threshold  
 438 varies from one model to another and it also depends on the large-scale parameterization  
 439 method used.



440 These relationships between  $CRH$  and mean precipitation rate are qualitatively similar  
 441 to that seen in observations over the tropical ocean regions [*Bretherton et al.*, 2004] (see  
 442 solid curves in Figure 5 ), and in other idealized models [e.g., *Raymond and Zeng*, 2005;  
 443 *Wang and Sobel*, 2011], but there are significant quantitative differences. For instance,  
 444 CRMs using either the WTG or DGW method produce similar relationship between  $P$   
 445 and  $CRH$ . However, all CRMs using either the WTG or DGW method have a higher  
 446 threshold than observations and their mean precipitation rates rise more abruptly with  
 447  $CRH$  than in observations (see Figures 5a and 5c). In contrast, SCMs show a much  
 448 larger variety of relationships (see Figures 5b and 5d). Moreover, the transition from  
 449 near zero precipitation to rapid increase in precipitation with  $CRH$  is sharper in some  
 450 models compared to others (e.g., compare  $P$  versus  $CRH$  in the WTG simulations of  
 451 UMv7.8 and LMDzB; stars and right facing triangles in Figure 5b, respectively). When  
 452  $P$  and  $CRH$  are scaled by their reference values (see Figure 6), the CRMs produce a  
 453 relatively tight relationship. The spread among SCMs is also clearly reduced, although  
 454 considerable scatter remains. In general, the threshold occurs at around  $CRH_{Ref}$  and  
 455 beyond that,  $P$  increases much more rapidly with  $CRH$  in CRMs than in SCMs.

### 4.3. Budget analysis

456 As in *Daleu et al.* [2015], we analyze the budgets in order to clarify the differences  
 457 among RCE, WTG, and DGW simulations. For a simulation with parameterized large-  
 458 scale circulation, the heat and moisture budgets are written as

$$459 \quad H + P + R + H_{LS} = 0 \quad \text{and} \quad E - P + M_{LS} = 0, \quad (7)$$

460 respectively.  $E$ ,  $H$ ,  $P$  and  $R$  denote the domain and time-averaged values of surface  
 461 evaporation, surface sensible heat flux, precipitation rate and vertically integrated radia-  
 462 tive cooling rate respectively. The heating rate and moistening rate due to the diagnosed  
 463 large-scale circulation ( $H_{LS} = C_p \langle \partial \bar{T} / \partial t \rangle_{LS}$  and  $M_{LS} = L_v \langle \partial \bar{q} / \partial t \rangle_{LS}$ , respectively) are  
 464 zero by definition for the RCE simulations.  $C_p$  is the heat capacity at constant pressure  
 465 and  $L_v$  is the latent heat of vaporization.

466 From the moisture budget equation, the changes in mean precipitation rate with respect  
 467 to the value of the RCE reference state,  $\Delta P$ , must be due to changes in surface evaporation  
 468 with respect to the value of the RCE reference state,  $\Delta E$ , and/or the moistening rate due  
 469 to the large-scale circulation  $M_{LS}$ . Figures 7 and 8 show scatter plots of  $\Delta P$  versus  $M_{LS}$   
 470 and scatter plots of  $\Delta P$  versus  $\Delta E$ , respectively.

471 Both CRMs and SCMs show fairly linear relationships between  $\Delta P$  and  $M_{LS}$ . However,  
 472 the slope is not one-to-one (dotted oblique line in Figure 7), which implies changes in  
 473 surface evaporation as shown in Figure 8.  $\Delta E$  increases with  $\Delta P$  in a large proportion  
 474 of the WTG and DGW simulations, and there are only a few simulations which show  
 475 an enhancement of convective activity associated with a reduction in surface evaporation  
 476 (e.g., WTG simulation of LaRC-CRM an SST of 300.5 K, dark green solid diamond in  
 477 Figure 8a) or which show a suppression in convective activity associated with an increase  
 478 in surface evaporation (e.g., the WTG simulation of EC-Earthv1 with SST of 300.5 K,  
 479 dark green diamond in Figure 8b).

480 The sensitivity of surface fluxes (sum of sensible heat and latent heat fluxes) to changes  
 481 in near-surface perturbation winds due to changes in convective activity has been some-  
 482 what constrained in this study by imposing a mean horizontal wind speed in the surface

483 flux calculations. As a result,  $\Delta E$  is generally much smaller than  $\Delta P$ , such that changes  
 484 in precipitation are largely balanced by the large-scale moistening rates. This is readily  
 485 seen in Figures 7 and 8. For a large proportion of the simulations, the values of  $M_{LS}$  are  
 486 about or more than two third the values of  $\Delta P$ .

487 We now examine the relationship between  $\Delta P$  and the normalized gross moist stability  
 488 (NGMS),  $\Gamma$ .  $\Gamma$  is defined as the dimensionless number which relates the net lateral outflow  
 489 of moist static energy from a convective region to a measure of the strength of convection  
 490 in that region [Raymond *et al.*, 2009]. That is

$$491 \quad \Gamma = -\langle \bar{\omega} \partial \bar{h} / \partial p \rangle / L_v \langle \bar{\omega} \partial \bar{q}_v / \partial p \rangle, \quad (8)$$

492 where  $h$  is the moist static energy. Following Daleu *et al.* [2015],

$$493 \quad \Gamma = -(M_{LS} + H_{LS}) / M_{LS}, \quad (9)$$

494 and a diagnostic equation for  $\Delta P$  is

$$495 \quad \Delta P = \frac{\Gamma + 1}{\Gamma} \Delta E + \frac{\Delta H + \Delta R}{\Gamma}, \quad (10)$$

496 where  $\Delta H$  and  $\Delta R$  are respectively the changes in surface sensible heat flux and column-  
 497 integrated radiative cooling rates with respect to the values of the RCE reference state.

498 The reader is referred to Daleu *et al.* [2015] for a derivation of equation 10.

499 As discussed above,  $\Delta H$  is much smaller than  $\Delta P$ .  $\Delta R$  is also much smaller than  $\Delta P$  as  
 500 a result of imposing a fixed radiative cooling profile throughout most of the troposphere.  
 501 Also, most of these simulations show that the sum of  $\Delta H$  and  $\Delta R$  is much smaller than  
 502  $\Delta E$ , such that the factor  $(\Gamma + 1)/\Gamma$  largely describes the strength of the relationship  
 503 between  $\Delta P$  and  $\Delta E$  (see equation 10).

504 For the WTG and DGW simulations which reproduce the RCE reference state to a  
 505 good approximation,  $\Gamma$  is a poor diagnostic since  $M_{LS} + H_{LS}$  and  $M_{LS}$  are both close to  
 506 zero, consistent with a weak large-scale circulation. Moreover,  $\Gamma$  measures the efficiency  
 507 of convection in removing moisture static energy from the column and thus, is not a  
 508 particularly useful diagnostic when convection is strongly suppressed. Therefore, the  
 509 values of  $\Gamma$  for the WTG and DGW simulations which result in significant large-scale  
 510 descent are not relevant, and we consider Figure 9, which shows  $\Gamma$  as a function of SST  
 511 for the WTG and DGW simulations which result in significant large-scale ascent only.  
 512 These are simulations which produce  $P/P_{Ref} > 1.1$  with  $\Omega > 0.4 \times 10^{-2} \text{ Pa s}^{-1}$ .

513 Most CRM simulations which result in significant large-scale ascent have positive values  
 514 of  $\Gamma$  and the WTG and DGW simulations of LaRC-CRM with an SST of 300.5 K are  
 515 the only CRM simulations which have negative values of  $\Gamma$  (black diamonds in Figures 9a  
 516 and 9c). Among SCMs, simulations with warm SSTs which produce significant large-scale  
 517 ascent have positive values of  $\Gamma$ . Negative  $\Gamma$  in some SCMs are obtained in the simulations  
 518 which result in either large-scale ascent over a cold SST (e.g., the WTG of GISS-SCM  
 519 with an SST of 299.5 K; Table 4 and black circles in Figure 9b) or large-scale ascent  
 520 over a uniform SST (e.g., the WTG and DGW simulations of EC-Earthv1 and the DGW  
 521 simulation of ARPV6 over a uniform SST of 300 K; Table 4 and red diamonds in Figures 9b  
 522 and 9d, and black down facing triangles in Figures 9d). In the simulations which result  
 523 in significant large-scale ascent and have negative values of  $\Gamma$ ,  $M_{LS}$  values are positive.  
 524 Therefore, negative values of  $\Gamma$  are the result of a deficit of cooling over moistening rates.  
 525 That implies a reduction in evaporation despite an increase in precipitation rate in those  
 526 simulations (e.g., dark green diamond in Figures 8a and c). With the exception of the

527 negative values of  $\Gamma$ ,  $\Gamma$  generally ranges between 0 and 1, with only few SCM simulations  
 528 having  $\Gamma > 1$  (e.g., LMDzA using the DGW with an SST of 300.5 K; blue left facing  
 529 triangles in Figure 9d).

530 CRMs (except LaRC-CRM) and three SCMs (EC-Earthv3, UMv7.8 and LMDzB) using  
 531 either the WTG or DGW method have  $\Gamma$  which is relatively insensitive to the SST. In  
 532 those models,  $\Delta E$ , and hence  $M_{LS}$  scale approximately linearly with  $\Delta P$ . In the other  
 533 four SCMs and LaRC-CRM,  $\Gamma$  show large sensitivity including non-monotonic behaviour,  
 534 and there are substantial differences in the relationship between  $\Gamma$  and SST depending on  
 535 which large-scale parameterization is used (e.g., compare  $\Gamma$  versus SST for the WTG and  
 536 DGW simulations of AR Pv6; down facing triangles in Figures 9b and 9d).

537 In this study, there is no straightforward relation between  $\Gamma$  and the top-heaviness of  
 538  $\bar{\omega}$  calculated as the mass-weighted vertical integral of the pressure velocity over the layer  
 539 at 500 – 100 hPa (see definition in Section 4.1). In addition,  $\Gamma$  does not explain the  
 540 difference between different models sensitivity to SST. Despite the fact that studies of  
 541 this nature allow convection to interact with the large-scale dynamics, there are many  
 542 differences between these feedbacks compared to full General Circulation Models (GCMs)  
 543 and the real tropical circulations. For instance, evaporation is not tied to the large-scale  
 544 circulation, the moisture convergence is directly tied to the dynamical convergence without  
 545 any contribution from the rotational part of the flow, and radiation is non interactive.  
 546 Therefore, the NGMS in these studies may have very different characteristics compared  
 547 to that of full GCMs and the real tropical circulations.

548 On the other hand, *Wang and Sobel* [2011] idealize horizontal moisture convergence and  
 549 radiation in the same way as in our study and found that  $\Gamma$  is a predictor of  $\Delta P$  in the

precipitating regime. In this study, only two models exhibit positive values of  $\Gamma$  which  
 are a monotonically decreasing function of SST or  $P$  as in *Wang and Sobel* [2011]. These  
 models are WRF and LEMv2.4 using either the WTG or DGW method (circles and down  
 facing triangles in Figures 9a and 9c). In contrast to the result of *Wang and Sobel* [2011],  
 some models exhibit positive values of  $\Gamma$  which are a monotonically increasing function of  
 SST or  $P$  (e.g., AR Pv6 using the WTG method with warm SSTs , down facing triangles  
 in Figure 9b) while other models exhibit positive or negative values of  $\Gamma$  which are not  
 directly related to SST or  $P$  (e.g., LaRC-CRM using the DGW method with warm SSTs,  
 diamonds in Figure 9c). In the latter case,  $\Gamma$  and  $\Delta E$  are both important to predict  $\Delta P$   
 (see equation 10).

## 5. Conclusions

In this international intercomparison project, we used the WTG and DGW methods to  
 study the two-way interaction between convection and large-scale circulations in various  
 CRMs and SCMs. Using the WTG method we derived the large-scale circulation that  
 reduces the virtual potential temperature anomalies over a given time-scale [*Raymond*  
*and Zeng*, 2005; *Sobel et al.*, 2007; *Sessions et al.*, 2010; *Daleu et al.*, 2012], and using  
 the DGW we simplified the large-scale circulation to a linear gravity wave of a single  
 horizontal wave number [*Kuang*, 2008, 2011; *Romps*, 2012a, b]. In both cases, the derived  
 large-scale circulation couples a model to a reference state defined with profiles generated  
 from previous RCE simulations of the same model. In *Daleu et al.* [2015], we analysed  
 WTG and DGW simulations over a uniform SST. In this paper, we kept the reference  
 state fixed and conducted WTG and DGW simulations with different values of SST in  
 the simulated column.

572 The WTG and DGW simulations with a cold (or a warm) SST result in lower (or higher)  
573 precipitation rates (compared to the value of the RCE reference state) in all CRMs and  
574 in a large proportion of the SCMs. In a few SCMs, a WTG simulation over a warm  
575 SST and a corresponding DGW simulation produce different signs of the circulation. In  
576 those SCMs, different signs of the circulation occur because the WTG simulation produces  
577 large-scale ascent over a cold SST or large-scale descent over a warm SST.

578 In general, the behavior across models for a given large-scale parameterization method  
579 is different, and the behavior of an individual model also depends on which large-scale  
580 parametrization is used. However, DGW simulations do produce large-scale pressure ve-  
581 locity profiles which are smoother than those produced by WTG simulations, and consis-  
582 tent with the results of *Wang et al.* [2013], DGW simulations generally produce large-scale  
583 pressure velocity profiles which are less top-heavy compared to those produced by WTG  
584 simulations.

585 All CRMs and five out of the seven SCMs show a monotonic increase of mean precipi-  
586 tation rate with SST using either the WTG or DGW method. A similar relationship  
587 between precipitation rate and SST was produced in *Sobel and Bretherton* [2000] and  
588 *Ramsay and Sobel* [2011]. The other two SCMs show sensitivity of the mean precipitation  
589 rate with SST which is not always monotonic. CRMs show a fairly linear relationship  
590 between mean precipitation rate and the amplitude of the diagnosed vertically-integrated  
591 large-scale circulation, while a few SCMs show deviations from this linear relationship,  
592 particularly for simulations with warm SST.

593 Precipitation is an increasing function of the column relative humidity, with the former  
594 increasing rapidly as the latter passes a threshold. A similar relationship is found in other

595 numerical modeling studies [*Wang and Sobel, 2011; Raymond and Zeng, 2005*], and is  
596 consistent with observations [*Bretherton et al., 2004; Holloway and Neelin, 2009*]. All  
597 CRMs using either the WTG or DGW method show a similar relationship between mean  
598 precipitation rate and column-relative humidity. They are all moister and the resulting  
599 mean precipitation rate increases more abruptly with column relative humidity than in  
600 observations. SCMs show a much wider range of relationships between precipitation rate  
601 and column-relative humidity, although this spread is reduced when values are normalized  
602 by their RCE values.

603 In our WTG and DGW simulations, the change in precipitation with respect to the  
604 value of the RCE reference column is largely balanced by the moistening rate due to the  
605 large-scale circulation. We calculated the NGMS for simulations with significant large-  
606 scale ascent at equilibrium. A large proportion of those simulations exhibited positive  
607 values of NGMS, ranging between 0 and 1, and only a few simulations exhibit negative  
608 values of NGMS or values of NGMS that approach 1.5. Those which exhibit negative  
609 values of NGMS have a deficit of cooling over moistening rates, which implies a reduction  
610 in evaporation despite an increase in precipitation rate. Most CRMs and three SCMs  
611 using either the WTG or DGW method show small sensitivity of the NGMS with the  
612 SST. In the other CRM and the other four SCMs the relationship between NGMS and  
613 SST varies considerably and depends on the large-scale parameterization method used.  
614 In this study  $\Gamma$  is not related to the shape of the large-scale pressure velocity profile  
615 and does not explain the difference between different model's sensitivity to SST. That  
616 is, in comparison to real tropical circulations, the NGMS in this configuration may not



617 be a very important diagnostic due to the way in which evaporation, horizontal moisture  
618 convergence and radiation are idealized.

619 In this intercomparison project convection feeds back on the large-scale forcing, the  
620 moisture source is induced by the derived large-scale motion, and the precipitation rate  
621 produced is the result of both the model physics and parametrized large-scale dynamical  
622 feedback. Therefore, this study can be viewed as an extension of traditional intercom-  
623 parisons with prescribed large-scale forcing (e.g., TOGA COARE and DYNAMO) and  
624 intercomparisons in which moisture source is defined as a relaxation to a prescribed pro-  
625 file [*Derbyshire et al.*, 2004]. The results from this intercomparison project are important  
626 for understanding the two-way interaction between convection and large-scale tropical  
627 dynamics and also for interpreting discrepancies between the results reported in the lit-  
628 erature. Our results suggested that the discrepancies between the published results can  
629 be related to the choice of the large-scale parameterization method. For instance, we  
630 found that an individual model can produce different equilibrium states depending on the  
631 large-scale parameterization method used.

632 Moreover, we found that even with exactly the same implementation of the WTG  
633 or DGW method, different SCM and even CRM models produce different sensitivities  
634 of the equilibrium state to SST. CRMs that participated in this study differ in their  
635 representation of subgrid scale processes that are important for the evolution of convection  
636 and its interaction with large-scale circulation (e.g., cloud microphysics) . The differences  
637 in CRMs lead to some diversity of behavior in RCE simulations [*Daleu et al.*, 2015],  
638 and the diversity of behavior can be amplified when the physics is allowed to interact  
639 with the large-scale dynamics. However, despite the diversity obtained among CRMs, our

640 study demonstrates much larger inter-model variability among SCMs. That is, despite the  
641 significant differences in CRMs (e.g., resolution, domain size, microphysics and etc), the  
642 behaviour of these simulations using models with explicit convection are more constrained  
643 than those with parameterized convection.

644 This study has evaluated CRM and SCM sensitivities to parameterized large-scale  
645 dynamical feedback with fixed radiation and a non interactive surface. Further study  
646 may compare models and large-scale parameterization methods with interactive radiation  
647 and/or an interactive surface. Since our study indicates that there is a greater consistency  
648 in the behavior of CRMs under parameterized large-scale circulation while SCMs produce  
649 a much larger variation of behaviors, comparison between CRM and SCM behavior under  
650 parameterized large-scale circulation may be a useful tool when developing and testing  
651 parameterization schemes. Therefore, further analysis may be to assess the impact of  
652 changes in parameterization within a particular SCM.

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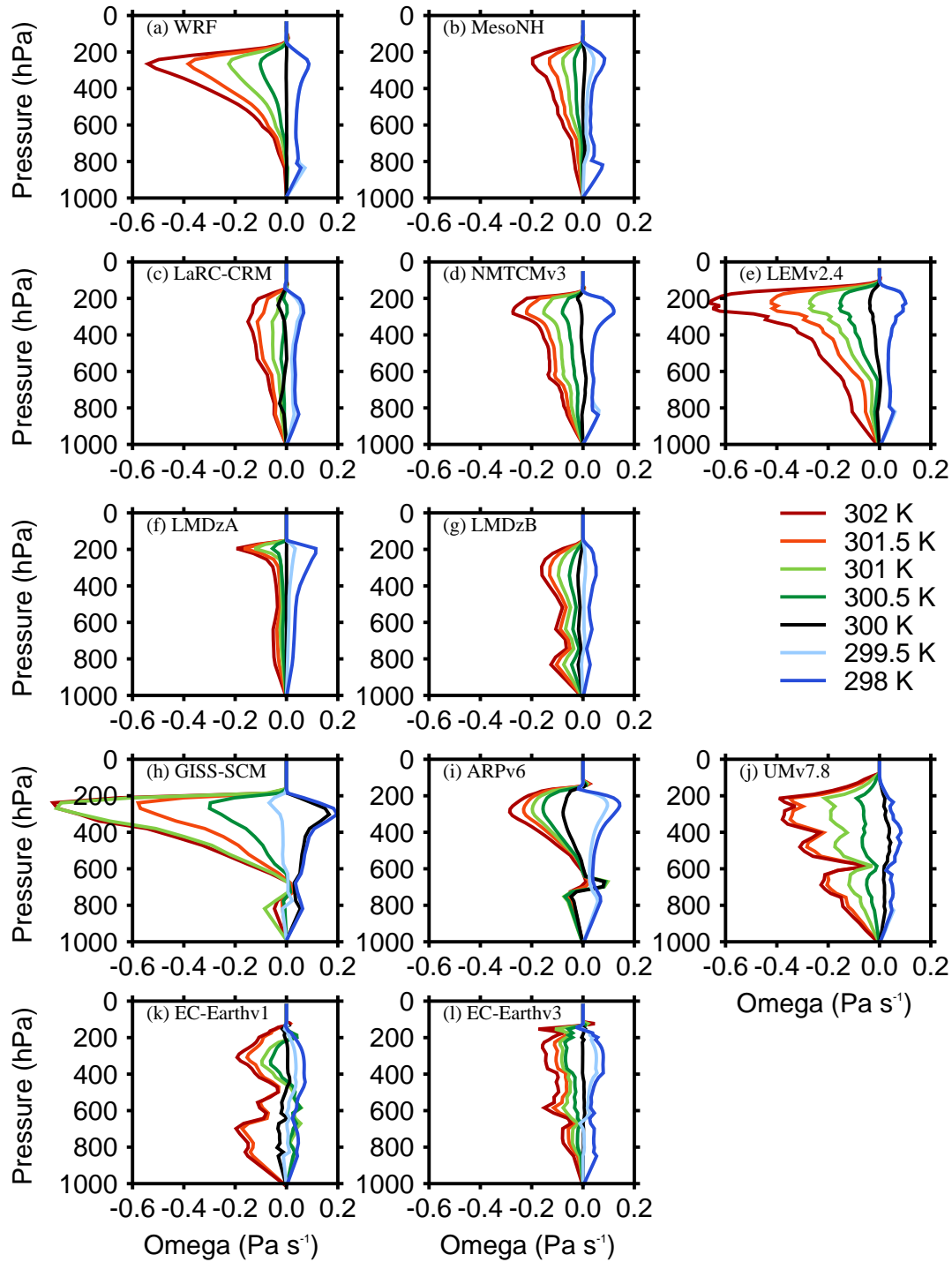


Model type	Cloud-Resolving Models (CRMs)				
Modelling group	Columbia University	CNRM-GAME	NASA	New Mexico Tech	UK Met Office
Model ID	WRF	MesoNH	LaRC-CRM	NMTCMv3	LEMv2.4
Symbol	●	▲	◆	■	▼
Dimension	3D	3D	2D	2D	2D
Hor. size (km)	190 × 190	150 × 150	256	200	128
Hor. res (km)	2 × 2	3 × 3	4	1	0.5
$P_{Ref}$ (mm d <sup>-1</sup> )	4.71	4.63	4.60	4.35	4.82

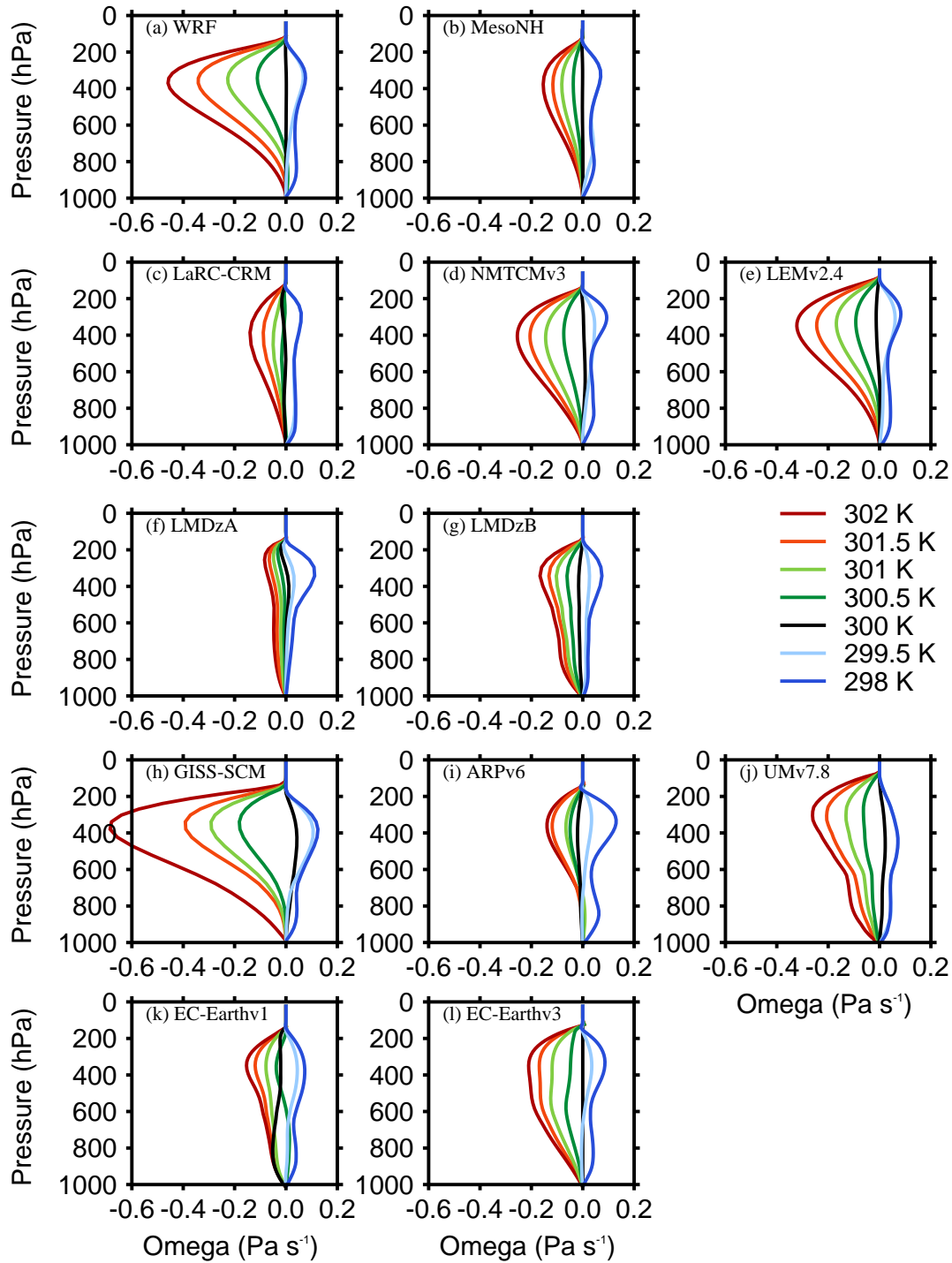
**Table 1.** List of cloud-resolving models (CRMs) that participated in this study. The symbols serve as a legend for results presented in Section 4.  $P_{Ref}$  is the mean precipitation rate obtained in the radiative-convective equilibrium simulation of each CRM with an SST of 300 K.

Model type	Single-Column Models (SCMs)						
Modelling group	LMD/IPSL		NASA	CNRM-GAME	UK Met Office	Koninklijk Nederlands Meteorologisch Insituut	
Model ID	LMDzA	LMDzB	GISS-SCM	ARPEGEv6 (ARPV6)	UMv7.8	EC-Earthv1	EC-Earthv3
Symbol	◁	▷	○	▽	★	◇	□
$P_{Ref}$ (mm d <sup>-1</sup> )	4.38	4.39	4.58	3.71	4.76	4.53	4.15

**Table 2.** *List of single-column models (SCMs) that participated in this study. The symbols serve as a legend for results presented in Section 4.  $P_{Ref}$  is the mean precipitation rate obtained in the radiative-convective equilibrium simulation of each SCM with an SST of 300 K.*



**Figure 1.** Large-scale pressure velocities obtained at equilibrium in the WTG simulations with an SST of 298 K (dark blue), 299.5 K (light blue), 300 K (black), 300.5 K (dark green), 301 K (light green), 301.5 K (orange), 302 K (red). Results are shown for the (a, b, c, d and e) CRMs and (f, g, h, i, j, k and l) SCMs. For each model, the reference profiles are their own RCE profiles at 300 K.



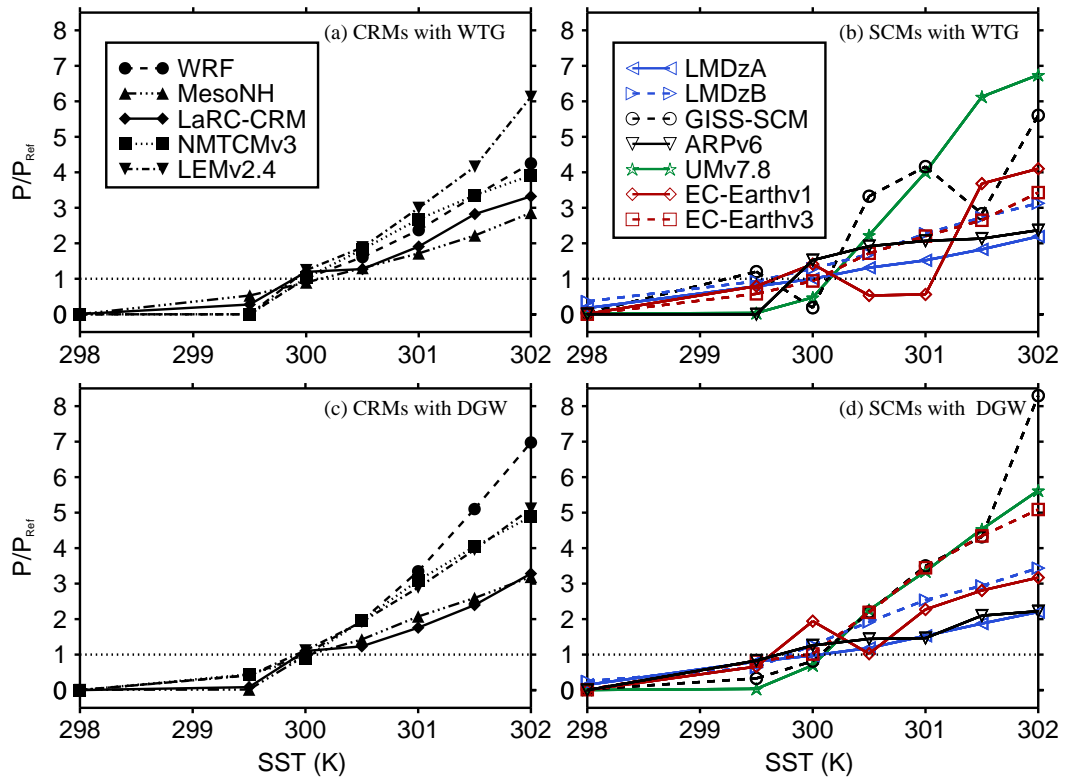
**Figure 2.** As in Figure 1, but for the equilibrium in the DGW simulations.

Model-CRMs	WTG or DGW	$P/P_{Ref}$ or $\Omega$	SST= 298 K	SST= 299.5 K	SST= 300 K	SST= 300.5 K	SST= 301 K	SST= 301.5 K	SST= 302 K
WRF	WTG	$P/P_{Ref}$	0.000	0.000	<b>1.020</b>	1.610	2.370	3.330	4.240
		$\Omega$	-4.210	-4.420	<b>0.180</b>	2.990	6.670	11.230	15.520
	DGW	$P/P_{Ref}$	0.000	0.420	<b>1.008</b>	1.950	3.350	5.100	6.970
		$\Omega$	-4.089	-2.575	<b>0.110</b>	4.180	9.485	15.547	22.235
MesoNH	WTG	$P/P_{Ref}$	0.002	0.529	0.896	1.303	1.714	2.220	2.857
		$\Omega$	-4.079	-1.577	<b>-0.290</b>	1.391	3.176	5.221	7.839
	DGW	$P/P_{Ref}$	0.009	0.015	<b>0.970</b>	1.425	2.073	2.594	3.190
		$\Omega$	-3.739	-3.461	<b>0.060</b>	1.803	4.145	6.0356	8.113
LaRC-CRM	WTG	$P/P_{Ref}$	0.006	0.276	1.200	1.233	1.908	2.824	3.322
		$\Omega$	-3.549	-2.621	0.970	0.887	2.994	6.111	7.856
	DGW	$P/P_{Ref}$	0.000	0.084	1.102	1.233	1.757	2.394	3.282
		$\Omega$	-3.563	-3.293	0.610	0.794	2.555	4.671	7.558
NMTCMv3	WTG	$P/P_{Ref}$	0.000	0.001	<b>1.028</b>	1.830	2.679	3.352	3.912
		$\Omega$	-4.517	-4.621	<b>0.100</b>	3.303	6.570	9.221	11.317
	DGW	$P/P_{Ref}$	0.000	0.445	0.896	1.954	3.090	4.044	4.887
		$\Omega$	-4.291	-2.266	<b>-0.388</b>	3.696	7.665	11.073	13.917
LEMv2.4	WTG	$P/P_{Ref}$	0.000	0.000	1.240	1.886	2.997	4.159	6.124
		$\Omega$	-4.588	-4.668	1.110	5.471	9.745	15.162	24.048
	DGW	$P/P_{Ref}$	0.000	0.413	1.117	1.923	2.888	3.953	5.111
		$\Omega$	-4.460	-2.658	0.464	4.129	8.103	12.436	17.031

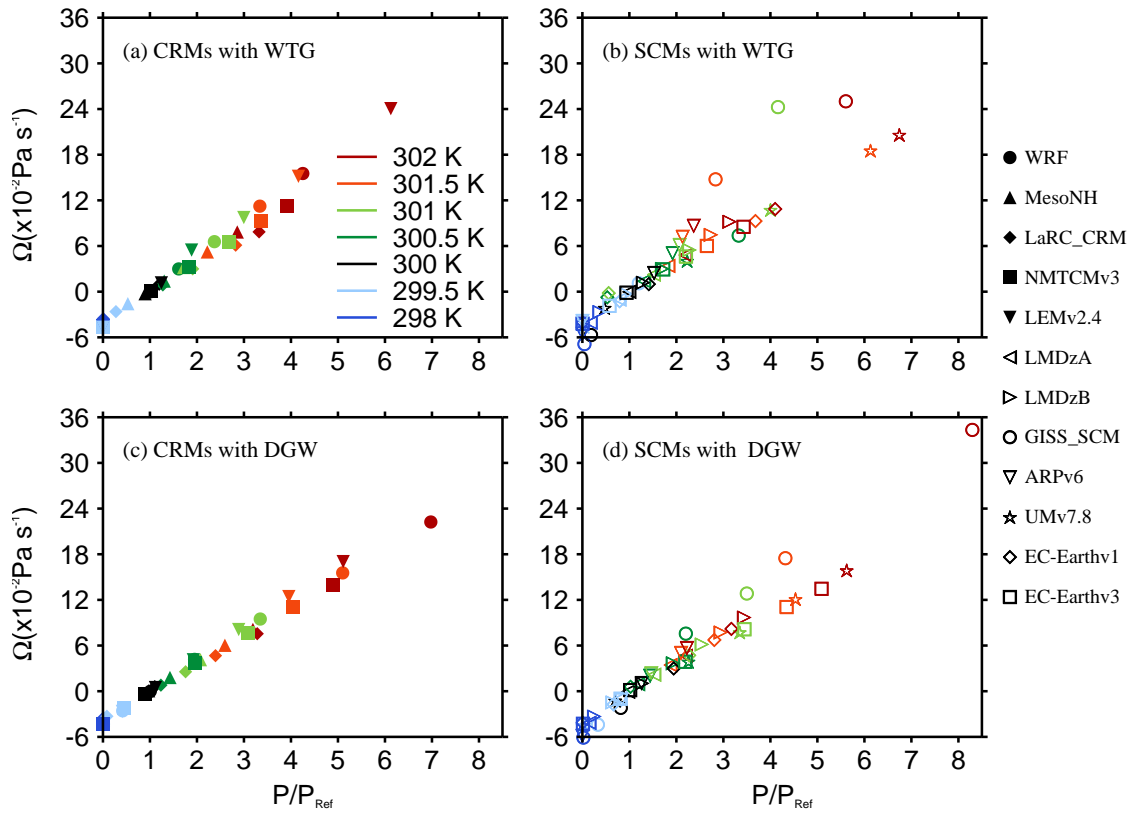
**Table 3.** Table showing the numerical values of  $\Omega$  ( $\times 10^{-2} Pa s^{-1}$ ) and  $P/P_{Ref}$  for WTG and DGW simulations with different values of SST in the simulated column. Results in bold correspond to  $|\Omega| < 0.4 \times 10^{-2} Pa s^{-1}$  (or  $\bar{\omega} \approx 0$ ) or  $0.9 < P/P_{Ref} < 1.1$ . If both  $\Omega$  and  $P/P_{Ref}$  are bold, the simulation with large-scale parameterization reproduces the RCE state to a good approximation.

Model-SCMs	WTG or DGW	$P/P_{Ref}$ or $\Omega$	SST= 298 K	SST= 299.5 K	SST= 300 K	SST= 300.5 K	SST= 301 K	SST= 301.5 K	SST= 302 K
LMDzA	WTG	$P/P_{Ref}$	0.176	0.790	<b>0.997</b>	1.313	1.520	1.829	2.192
		$\Omega$	-4.071	-1.037	<b>-0.015</b>	1.385	2.240	3.387	4.806
	DGW	$P/P_{Ref}$	0.15	0.804	<b>0.982</b>	1.187	1.530	1.874	2.201
		$\Omega$	-4.145	-0.972	<b>-0.065</b>	0.931	2.169	3.437	4.652
LMDzB	WTG	$P/P_{Ref}$	0.362	0.929	1.290	1.694	2.273	2.729	3.127
		$\Omega$	-2.670	-0.30	1.180	2.992	5.475	7.470	9.193
	DGW	$P/P_{Ref}$	0.248	0.638	1.269	1.922	2.537	2.940	3.437
		$\Omega$	-3.325	-1.462	1.030	3.676	6.153	7.726	9.689
GISS-SCM	WTG	$P/P_{Ref}$	0.044	1.200	0.180	3.325	4.161	2.833	5.605
		$\Omega$	-6.888	1.100	-5.700	7.371	24.25	14.760	25.022
	DGW	$P/P_{Ref}$	0.021	0.330	0.820	2.201	3.498	4.319	8.296
		$\Omega$	-6.095	-4.395	-2.180	7.566	12.837	17.475	34.335
ARPV6	WTG	$P/P_{Ref}$	0.003	0.000	1.530	1.920	2.067	2.132	2.368
		$\Omega$	-5.486	-3.852	2.230	5.055	6.132	7.210	8.658
	DGW	$P/P_{Ref}$	0.000	0.832	1.260	1.442	1.464	2.098	2.223
		$\Omega$	-5.853	-1.122	0.972	2.046	2.340	5.0256	5.673
UMv7.8	WTG	$P/P_{Ref}$	0.022	0.036	0.470	2.228	4.002	6.129	6.743
		$\Omega$	-4.528	-4.600	-2.130	4.053	10.751	18.537	20.610
	DGW	$P/P_{Ref}$	0.003	0.0343	0.700	2.257	3.350	4.534	5.623
		$\Omega$	-4.465	-4.437	-1.240	3.875	7.734	12.104	15.878
EC-Earthv1	WTG	$P/P_{Ref}$	0.011	0.792	1.420	0.529	0.558	3.682	4.101
		$\Omega$	-4.060	-1.262	0.990	-0.741	-0.192	9.275	10.855
	DGW	$P/P_{Ref}$	0.002	0.662	1.920	1.024	2.271	2.807	3.170
		$\Omega$	-4.117	-1.737	2.990	0.583	4.713	6.736	8.187
EC-Earthv3	WTG	$P/P_{Ref}$	0.003	0.577	<b>0.940</b>	1.720	2.202	2.648	3.430
		$\Omega$	-4.209	-1.860	<b>-0.135</b>	2.927	4.611	6.008	8.523
	DGW	$P/P_{Ref}$	0.0122	0.813	<b>1.014</b>	2.191	3.448	4.344	5.087
		$\Omega$	-4.280	-0.986	<b>0.146</b>	3.873	8.138	11.061	13.460

**Table 4.** Same as Table 3, but lists SCM results.

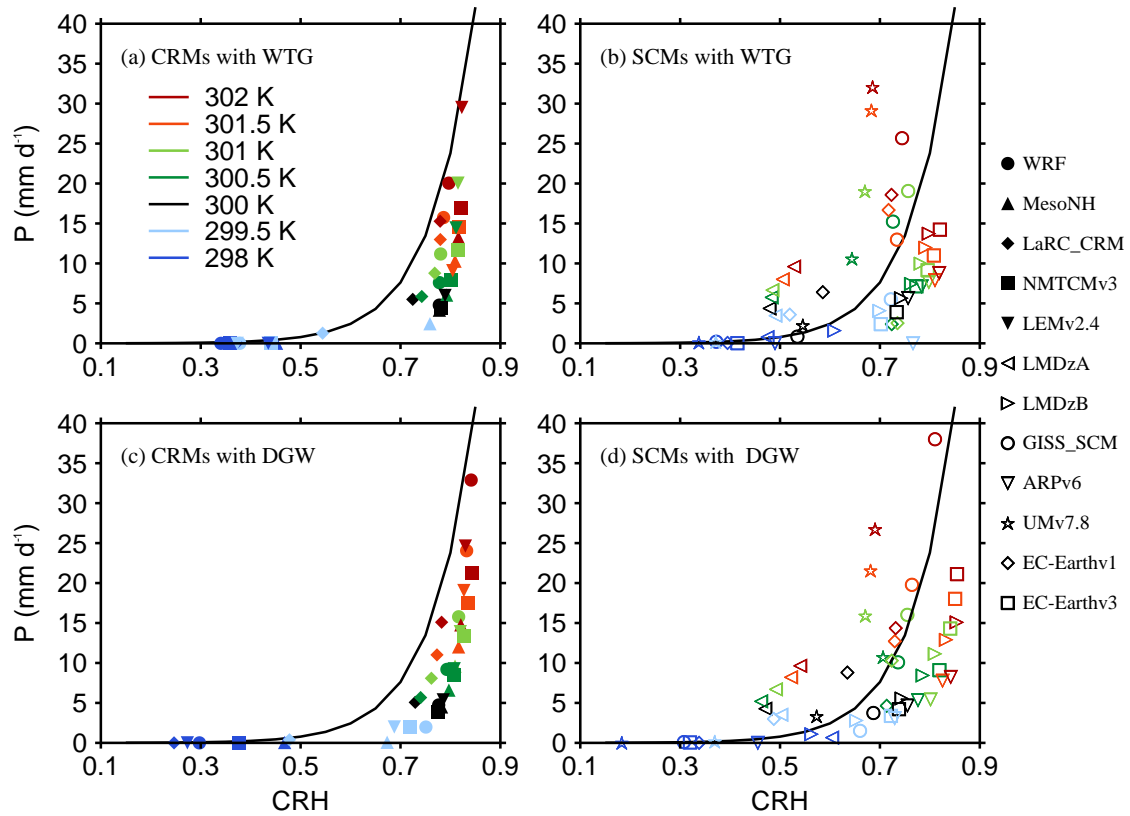


**Figure 3.**  $P/P_{Ref}$  versus SST. The values of  $P$  are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations. Results are shown for (left) CRMs and (right) SCMs. Symbol definitions are as in Tables 1 and 2.

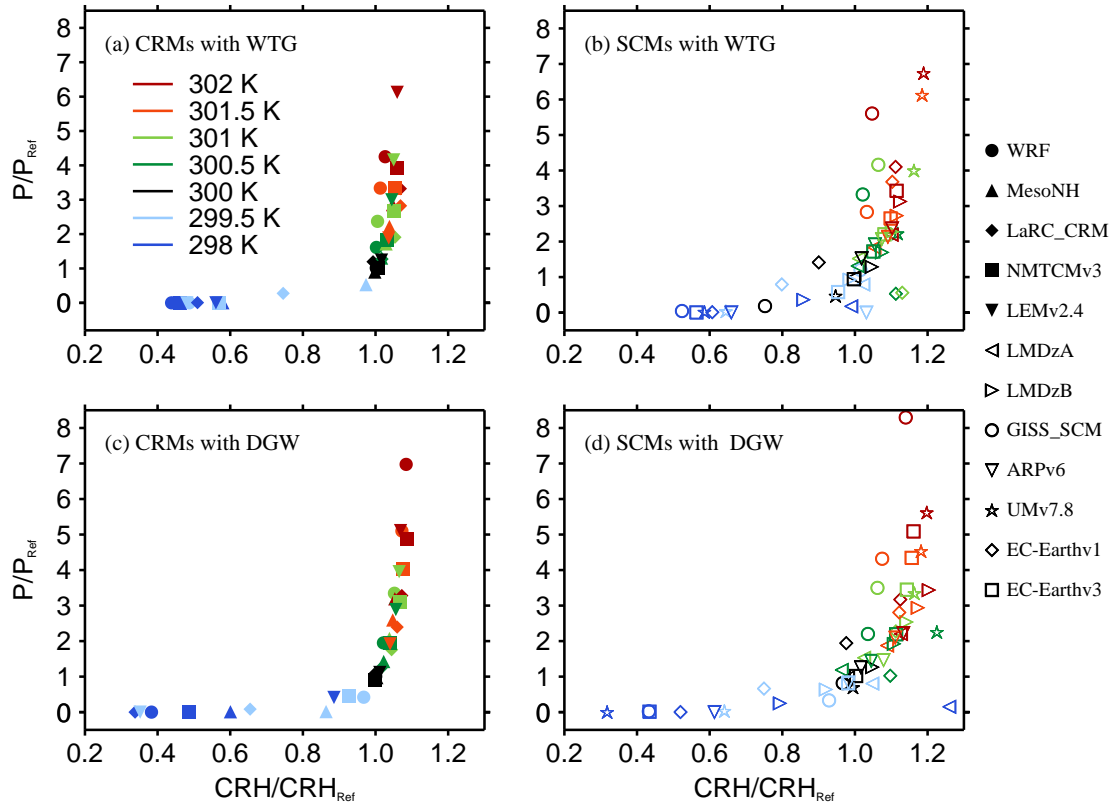


**Figure 4.** Scatter plots of  $\Omega$  versus  $P/P_{\text{Ref}}$ . Results are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations with an SST of 298 K (dark blue), 299.5 K (light blue), 300 K (black), 300.5 K (dark green), 301 K (light green), 301.5 K (orange), and 302 K (red). Results are shown for (left) CRMs and (right) SCMs. Symbol definitions are as in Tables 1 and 2.

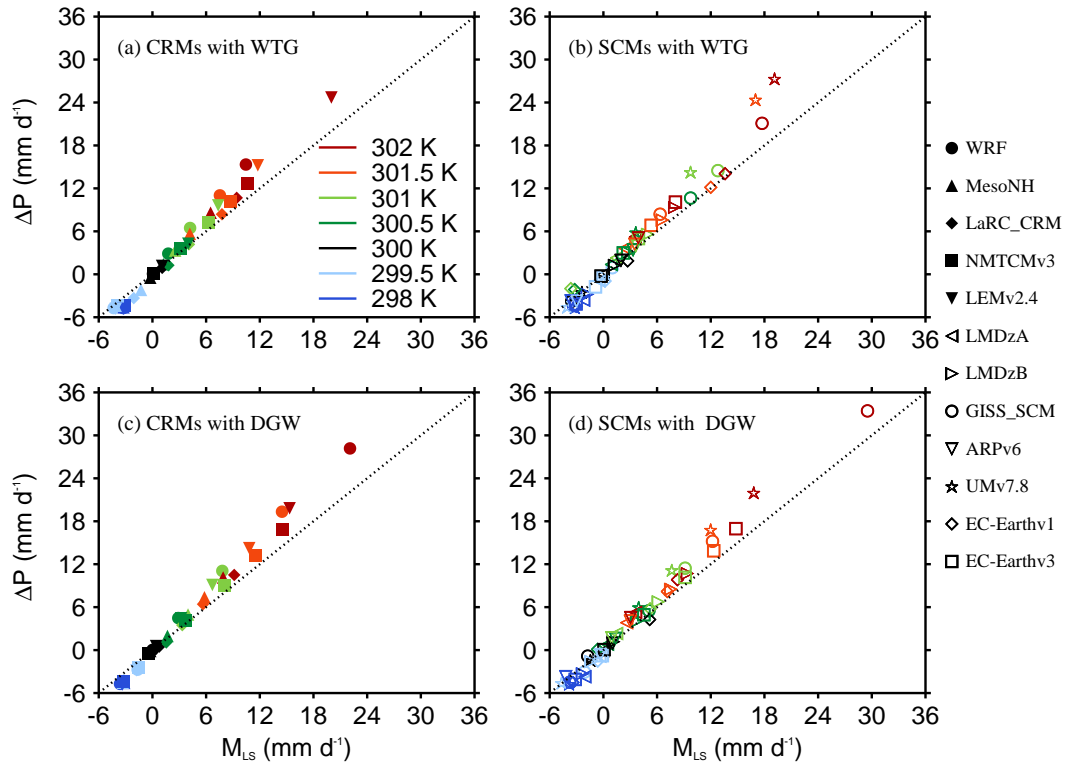




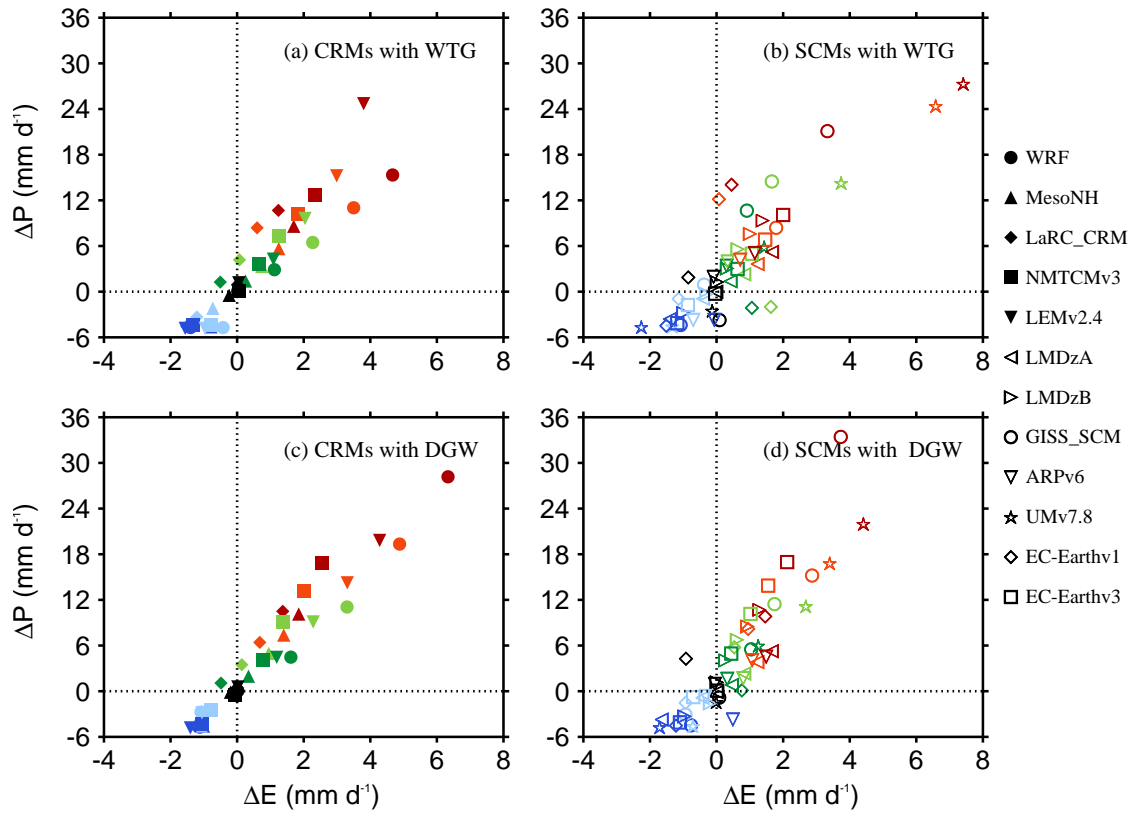
**Figure 5.** Scatter plots of  $P$  versus CRH (column relative humidity; the column-integrated water vapor divided by its saturation value). The results are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations with an SST of 298 K (dark blue), 299.5 K (light blue), 300 K (black), 300.5 K (dark green), 301 K (light green), 301.5 K (orange), and 302 K (red). Results are shown for (left) CRMs and (right) SCMs. Symbol definitions are as in Tables 1 and 2. The solid curve is the exponential fit for the observed monthly mean precipitation over the tropical oceans obtained by Bretherton et al. [2004].



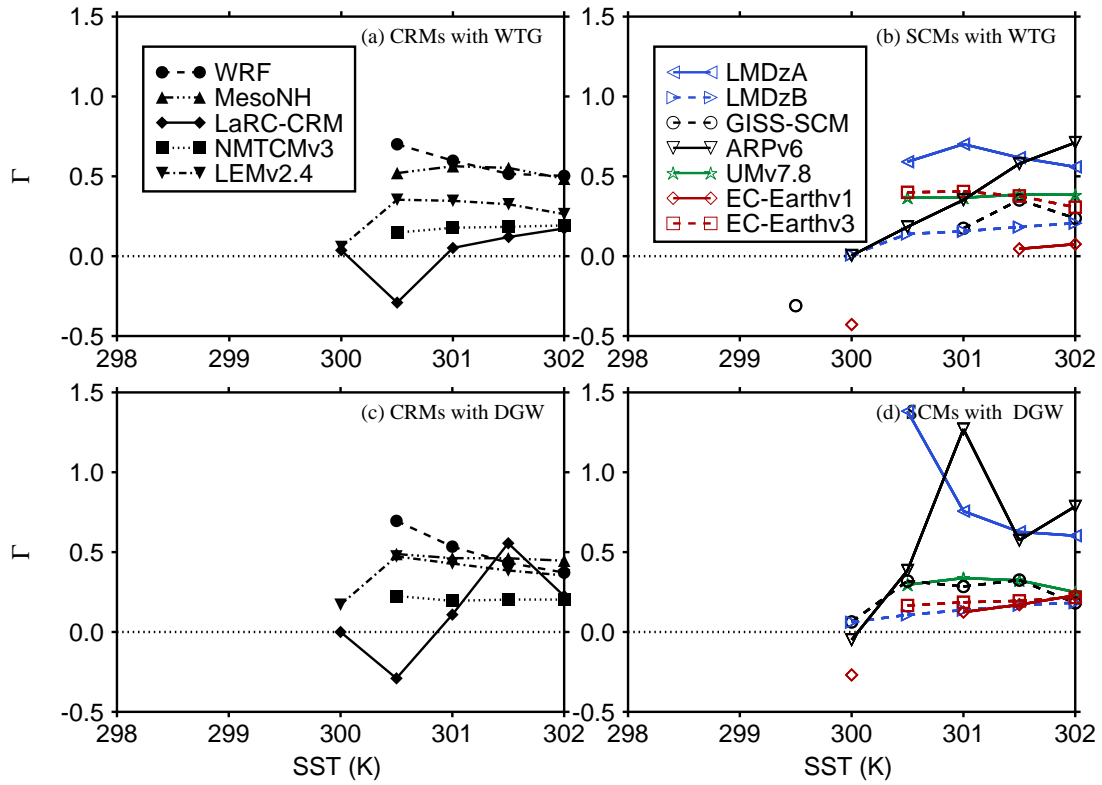
**Figure 6.** Scatter plots of  $P/P_{Ref}$  versus  $CRH/CRH_{Ref}$ , where  $CRH_{Ref}$  is the column relative humidity of the corresponding RCE reference state. The results are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations with an SST of 298 K (dark blue), 299.5 K (light blue), 300 K (black), 300.5 K (dark green), 301 K (light green), 301.5 K (orange), and 302 K (red). Results are shown for (left) CRMs and (right) SCMs. Symbol definitions are as in Tables 1 and 2.



**Figure 7.** Scatter plots of  $\Delta P$  versus  $M_{LS}$ . The results are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations with an SST of 298 K (dark blue), 299.5 K (light blue), 300 K (black), 300.5 K (dark green), 301 K (light green), 301.5 K (orange), and 302 K (red). Results are shown for (left) CRMs and (right) SCMs. The dotted oblique line corresponds to  $\Delta P = M_{LS}$ . Symbol definitions are as in Tables 1 and 2.



**Figure 8.** Scatter plots of  $\Delta P$  versus  $\Delta E$ . Results are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations over an SST of 298 K (dark blue), 299.5 K (light blue), 300 K (black), 300.5 K (dark green), 301 K (light green), 301.5 K (orange), and 302 K (red). Results are shown for (left) CRMs and (right) SCMs. Symbol definitions are as in Tables 1 and 2.



**Figure 9.**  $\Gamma$  versus SST. The values of  $\Gamma$  are those obtained at equilibrium in the (top) WTG and (bottom) DGW simulations which produce significant large-scale ascent only ( $P/P_{Ref} > 1.1$  with  $\Omega > 0.4 \times 10^{-2} \text{ Pa s}^{-1}$ ). Results are shown for (left) CRMs and (right) SCMs. Symbol definitions are as in Tables 1 and 2.