

# LongRunMIP - motivation and design for a large collection of millennial-length AO-GCM simulations

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### ABSTRACT

We present a model intercomparison project, LongRunMIP, the first collec-62 tion of millennial-length (1000+ year) simulations of complex coupled cli-63 mate models with a representation of ocean, atmosphere, sea ice, and land 64 surface, and their interactions. Standard model simulations are generally only 65 a few hundred years long. However, modeling the long-term equilibration 66 in response to radiative forcing perturbation is important for understanding 67 many climate phenomena, such as the evolution of ocean circulation, time-68 and temperature-dependent feedbacks, and the differentiation of forced signal 69 and internal variability. The aim of LongRunMIP is to facilitate research into 70 these questions by serving as an archive for simulations that capture as much 71 of this equilibration as possible. The only requirement to participate in Lon-72 gRunMIP is to contribute a simulation with elevated, constant  $CO_2$  forcing 73 that lasts at least 1000 years. LongRunMIP is a MIP of opportunity in that 74 the simulations were mostly performed prior to the conception of the archive 75 without an agreed-upon set of experiments. For most models, the archive 76 contains a preindustrial control simulation and simulations with an idealized 77 (typically abrupt)  $CO_2$  forcing. We collect 2D surface and top-of-atmosphere 78 fields, and 3D ocean temperature and salinity fields. Here, we document the 79 collection of simulations and discuss initial results, including the evolution of 80 surface and deep ocean temperature and cloud radiative effects. As of sum-81 mer 2019, the collection includes 50 simulations of 15 models by 10 modeling 82 centers. The data of LongRunMIP are publicly available. We encourage sub-83 mission of more simulations in the future. 84

(Capsule Summary) LongRunMIP is the first collection of millennial-length simulations of com plex coupled climate models and enables investigations of how these models equilibrate in re sponse to radiative perturbations.

### **1.** Motivation and objectives

Millennial-length climate simulations are necessary to understand the equilibrium states that occur in response to external forcings, as well as the relationship between transient and equilibrated behavior. Unforced millennial-length simulations are useful as well, as they allow us to consider long-term internal variability and to analyze shorter-term variability with increased statistical certainty. Reasons to study these long time scales include:

To better understand long-term climate dynamics. Outstanding issues include the time scales
 of ocean circulation response (e.g., Jansen et al. 2018; Rind et al. 2018), continental drying
 trends (e.g., Sniderman et al. 2019), or sea level rise (e.g., Bilbao et al. 2015; Rugenstein et al.
 2016c).

To help predict the impacts of 20th and 21st century emissions on century timescales, such as
 ice sheet stability, deep ocean warming, or polar amplification (e.g., Frölicher and Joos 2010;
 Clark et al. 2016; Mauritsen and Pincus 2017), which are rarely explicitly simulated using a
 fully-coupled climate model.

• To more accurately estimate Equilibrium Climate Sensitivity (ECS), which is the equilibrium response of the surface air temperature to a doubling of  $CO_2$  due to the "fast" feedbacks water vapor, lapse rate, clouds, and sea ice but excluding Earth system feedbacks such as changes in the carbon cycle, ice sheets, or vegetation. While ECS has long been a focus of scientific 106

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inquiry, substantial uncertainty remains as to its value (e.g., Charney et al. 1979; Knutti et al. 2017).

108	To understand the relationship between the transient response of the climate and its equilibra-
109	tion. Since radiative feedbacks can depend on the evolution of the spatial pattern of warming
110	(e.g., Senior and Mitchell 2000; Winton et al. 2010; Armour et al. 2013; Andrews et al. 2015;
111	Andrews and Webb 2018) and on the background temperature (e.g., Colman and McAvaney
112	2009; Caballero and Huber 2013; Block and Mauritsen 2013; Meraner et al. 2013; Bloch-
113	Johnson et al. 2015), a constant effective sensitivity of the climate is an inadequate assump-
114	tion. Several methods have been proposed to predict the equilibrium response from transient
115	simulations given a changing global feedback (Held et al. 2010; Winton et al. 2010; Armour
116	et al. 2013; Geoffroy et al. 2013b,a; Frölicher et al. 2014; Proistosescu and Huybers 2017;
117	Saint-Martin et al. 2019), but only fully equilibrated climate model simulations can serve to
118	test how well these methods predict equilibrium conditions.
119	To test theories for the relationship between feedbacks at different time-scales (Gregory et al.
120	2015, 2016; Zhou et al. 2016; Rugenstein et al. 2016a; Armour 2017; Proistosescu and Huy-
121	bers 2017; Ceppi and Gregory 2017; Andrews and Webb 2018; Andrews et al. 2018), and
122	to quantify the influence of slow, centennial-scale modes on the temperature evolution of the
123	last century (Armour 2017; Proistosescu and Huybers 2017).
124	To understand the relevance, time scales, and magnitude of the energy imbalances and drifts
125	exhibited by climate models (e.g., Hobbs et al. 2016), with the potential application of de-

• To understand the relationship between the forced response and internal variability. This relationship is currently studied using the time frame of one or two centuries, which is not

creasing the spin-up time needed to run these models.

129	enough to robustly quantify the internal variability under consideration (e.g., Maher et al.
130	2018; Lutsko and Takahashi 2018; Bloch-Johnson et al. in revision), millennial time scales
131	with varying forcings (e.g., Köhler et al. 2017; Khon et al. 2018; Rehfeld et al. 2018) or by
132	using expensive large ensemble simulations on decadal to centennial time scales (e.g., Deser
133	et al. 2012; Maher et al. 2019; Rodgers et al. 2015). Millennial long simulations allow us to
134	differentiate the transient response from the equilibrated forced response, even for quantities
135	with large internal variability, such as precipitation, droughts, or the El Niño-Southern Os-
136	cillation (ENSO), and also the significance of a change in internal variability in a transient
137	simulation relative to the control simulation (e.g., Brown et al. 2017).
138	• To compare climate model responses and paleo proxies, e.g. of surface or deep ocean temper-
139	atures or hydrological conditions on land in order to provide an independent way of testing
140	climate models (Gebbie and Huybers 2019; Burls and Fedorov 2017; Scheff et al. 2017).
141	With LongRunMIP, we aim to advance knowledge in the above mentioned areas, fill a gap in the
142	CMIP protocols (Taylor et al. 2011; Eyring et al. 2016), and collect published data in one location
143	for easy public access.
144	The goals of LongRunMIP are
145	a) to continuously gather existing millennial-length simulations (both published and unpub-
146	lished)
147	b) to standardize the collected data (e.g., using the same units and sign conventions)
148	c) to make the data publicly available and easily accessible
149	d) to foster an interdisciplinary community of users working on millennial-length problems,
150	with experts on oceanography, atmospheric dynamics, energy balance modeling, ice sheet
151	modeling, and paleoclimatology

### <sup>152</sup> The objectives of this paper are to

- a) motivate the data collection strategy (Section 2)
- b) specify the requirements for LongRunMIP contributors (Section 2 and b)
- c) give an overview of currently submitted simulations and models (Section 2a, b, and Table 1)
- d) give a sample of some initial analysis on these simulations (Section 3)
- e) show how LongRunMIP relates to the existing literature on millennial-length simulations
   (Section 4a)
- f) discuss the limitations and opportunities of LongRunMIP (Section 4b and c).

### **2. Experimental design and data collection strategy**

LongRunMIP is the first and largest compilation of millennial-length simulations of complex cli-161 mate models to date, where a "complex climate model" is understood to include an atmospheric, 162 sea ice, land, and full depth ocean component, i.e. Atmosphere-Ocean General Circulation Mod-163 els (AO-GCMs) with a dynamic atmosphere and ocean, as opposed to Models of Intermediate 164 Complexity (EMICs), which are often used to study millennial-length questions in climate science 165 (e.g., Zickfeld et al. 2013; Levermann et al. 2013). These model simulations include the "fast" 166 feedbacks, such as changes in water vapor, lapse rate, sea ice, and clouds (Charney et al. 1979), 167 but no "slow" feedbacks, such as changes in the ice-sheets. Vegetation is treated differently in the 168 models (see Section 2b). In Section 4 we discuss the implications and limitations of our approach. 169 Our goal is to collect as many simulations from as many independent models as possible, while 170 keeping the archive and data sharing manageable. Consequently, we keep our requirements for 171 contributions low. 172

A step-increase in atmospheric  $CO_2$  concentrations (in the following called "step-forcing") is 174 one of the simplest experiments for studying a model's response to forcing and is used as a bench-175 mark simulation in CMIP3, CMIP5, and CMIP6 (Meehl et al. 2007; Taylor et al. 2011; Eyring 176 et al. 2016). More realistic, gradual forcing scenarios have been shown to be representable by the 177 step-forcing scenarios and exhibit feedbacks that correlate with those computed from step-forcing 178 simulations (Good et al. 2013, 2015; Geoffroy and Saint-Martin 2014; Colman and Hanson 2016). 179 The CMIP3 protocol required a step-forcing of *doubling* atmospheric CO<sub>2</sub> (here referred to as 180 abrupt2x) above pre-industrial levels in a slab (i.e. non-dynamical) ocean, which for decades has 181 been used to define ECS (e.g., Charney et al. 1979; Boer and Yu 2003c; Danabasoglu and Gent 182 2009). The integration time scale of these model setups are a couple of decades. However, a 183 quadrupling of CO<sub>2</sub> (here referred to as *abrupt4x*) above pre-industrial levels has a better ratio of 184 forced signal to internal variability. Because the forced response was assumed to scale linearly 185 with increased forcing, the CMIP5 protocol requested an abrupt quadrupling of CO<sub>2</sub>, now in a 186 fully coupled model with a dynamical ocean, requiring longer integration time scales. The CMIP6 187 protocol again requests abrupt  $CO_2$  quadrupling experiments, but encourages also the submission 188 of abrupt  $CO_2$  doublings, to study the relation between different forcing levels (Eyring et al. 2016; 189 Good et al. 2016). CMIP5 and 6 protocols require the submission of 150 years of model output. 190 A representative response of surface temperature anomalies and top of the atmosphere (TOA) ra-191 diative imbalance to an *abrupt4x* scenario is shown in Fig. 1. All anomalies mentioned in this 192 paper are computed as the difference of the experiment from the average of the control simulation. 193 After the 150 years of CMIP protocol length (blue shading) and after 1000 years (the minimum 194 contribution to LongRunMIP, light red shading), the surface temperature response of the exem-195

<sup>196</sup> plary model shown here has reached 75 % and 88 % of its final value respectively, while the TOA <sup>197</sup> radiation has equilibrated 85 % and 93 % of the forcing respectively (7.6 W m<sup>-2</sup> for this model). <sup>198</sup> Thus, the final equilibration is a CPU-intensive exercise; the model shown here needs 4000 years <sup>199</sup> to balance the final 0.5 W m<sup>-2</sup> (dark red shading).

The set of variables we collect is motivated by the interest of the LongRunMIP contributors and 200 organizers in ECS, temperature and time dependent feedbacks, and deep ocean warming. Table 201 1 lists the variable names, units, and temporal and spatial resolution of the requested variables. 202 The naming and sign conventions follow the CMIP5 protocol<sup>1</sup>. Given the large amount of data 203 involved, we have kept our requested variable list low to allow as many groups as possible to 204 participate. For the same reason, we do not request the data to be "CMORized"<sup>2</sup>, i.e. written in 205 conformance with the CMIP standards. However, we do homogenize signs, variable long names, 206 and units, and also provide a regridded version of the fields, as well as global means. 207

### <sup>208</sup> b. Minimal, optimal, and current contributions

The *minimal requirement* to contribute to LongRunMIP are annual fields of a single simulation of any CO<sub>2</sub> forcing scenario that has at least 1000 years of constant forcing, along with a control simulation of any length. The complexity of the model should be CMIP5-class and include dynamic atmosphere, ocean, and sea ice components. An *optimal contribution* comprises monthly fields of fully equilibrated *abrupt2x*, 4x, and 8x simulations and a control simulation of several millennia.

Table 2 lists the model characteristics of the current contributions. Because the archive is assembled from experiments initiated independently for research purposes by multiple modeling groups, there is no pre-defined protocol like for the CMIP simulations. The models are diverse in origin

<sup>&</sup>lt;sup>1</sup>http://cmip-pcmdi.llnl.gov/cmip5/data\_description.html
<sup>2</sup>https://pcmdi.llnl.gov/CMIP6/Guide/dataUsers.html

and sample the CMIP5 range of models well (see discussion on model genealogy in Knutti 2010).
Table 2 lists references for each model and publications using (parts of) the model output. Most
of the current contributions to LongRunMIP are extensions of CMIP5 simulations, sometimes
with updated model versions, while one model is an extension of a CMIP3 and another model an
extension of a CMIP6 contributions (CCSM3 and CNRM-CM6-1 respectively).

Many of our current contributions fall short of the optimal expectation for equilibrium, because even several millennia are insufficient for the deep ocean to equilibrate (see discussion around Fig. 4). However, a few millennia appear to be enough for the surface temperature and TOA radiative imbalance to reach a new steady state in most models (see Section 3), and many questions can be adequately addressed with the current contributions. Our approach is to be inclusive, and to leave it to the user to determine the degree of equilibration needed for their research and to develop criteria for model selection.

Most contributions are step-forcing simulations, generally to 2x or 4x pre-industrial CO<sub>2</sub> con-230 centrations (in Fig. 2 *abrupt2x* in colored in yellow, *abrupt4x* in orange, *abrupt8x* in dark red; 231 *abrupt2.4x* and *abrupt4.8x* in dark and light pink). There are currently three exceptions: 1) some 232 model simulations have gradual increases in CO<sub>2</sub> at 1% per year until doubled or quadrupled con-233 centrations are reached, after which the concentration is kept constant (1pct2x and 1pct4x, light)234 and medium red in Fig. 2). 2) One model simulates the 1850-2010 period, after which  $CO_2$  in-235 creases either piecewise linearly for 90 years until reaching 2.4x pre-industrial values (CCSM3II). 236 3) Finally, one model simulates the historical period and then the CMIP5 extended representative 237 concentration pathway 8.5 (including CH<sub>4</sub>,  $N_2O$ , CFC11, and CFC12 in addition to  $CO_2$ ) until 238 year 2300 after which all forcing agents are kept constant (RCP8.5+, violet in Fig.2) For the 239 models that did not contribute a a millennial-long step-forcing simulation, we collect short (typ-240 ically 150 year) step-forcing simulations, generally from the CMIP5 archive. These simulations 241

can be used to estimate the effective climate sensitivity and to relate transient and equilibrium
responses. They are not mentioned in Table 2 and Fig. 2.

Most contributors were able to submit all requested variables. Some models only stored annual output, while for a few models the entire model output (including many more variables than listed in Table 1) is available. In principle, but with considerable effort, additional variables not listed in Table 1 could be requested from some or all contributors.

Some models are outliers in some sense. For example, the simulation *abrupt4x* of FAMOUS 248 warms anomalously strong (Fig. 2 and 7) due to a shortwave cloud effect which is positive through-249 out the simulation and longwave clear-sky effect, which increases anomalously strongly (not 250 shown, see Rugenstein et al. (2019)). In principle though, such extreme behavior could represent 251 possible characteristics of the real world (e.g., Bloch-Johnson et al. 2015; Schneider et al. 2019). 252 Another atypical model is EC-Earth-PISM, which is the only model with an interactive Greenland 253 ice sheet. This additional component and its historical and RCP8.5+ forcing scenario makes it 254 harder to compare the simulation to other models and attribute changes to one forcing component. 255 This model also does not equilibrate but finally produces a negative TOA imbalance, which prob-256 ably would increase if the simulation was integrated further. We encourage similar "problematic" 257 submissions, since our focus is on understanding model behavior and the large range of model 258 responses (discussed in Section 3). 259

In nine models, the vegetation is fixed to pre-industrial conditions (ECHAM5, CCSM3, CCSM3II, HadCM3L, FAMOUS, MIROC32, ECEARTH, GISSE2R, CNRMCM61), while the other seven models have dynamic vegetation schemes (MPIESM11, MPIESM12, CESM104, HadGEM2, GFDLESM2M, GFDLCM3, IPSLCM5A).

### **3. Sample of model output**

### *a. Imbalances in the control simulation and drift*

In principle, the TOA radiative imbalance should be zero in a control simulation. Most models 266 contributing to LongRunMIP do not loose or gain energy (Fig. 3). However, some models that are 267 equilibrated in the sense that they show no substantial drift, still have a constant energy leakage. 268 For CMIP5 models, imbalances of the same order of magnitude (and larger) have been shown to be 269 uncorrelated with the forced response (Hobbs et al. 2016). If computing atmospheric anomalies, 270 we suggest users to take the difference of each time step to the time-averaged control simulation 271 imbalance, except for CCSM3II and GFDL-CM3 for which the difference to a polynomial fit to 272 the control simulation time series seems appropriate (see Fig. 3). 273

The deep ocean (defined here as depth level around 2 km) has an astonishingly small drift in 274 the global average in most models (Fig. 4, lowest panel). While the surface ocean time scales 275 closely follows the global mean surface air temperature anomaly, the deep ocean takes centuries 276 to equilibrate. Panel a and b of Fig. 4 display the surface and deep ocean temperature anomalies, 277 computed as the difference of the forced and control simulations, while the lowest panel shows the 278 absolute temperatures of the deep ocean in the control simulations to indicate the model spread in 279 the base state. Previous work on long-term trends in deep ocean temperature and salinity shows 280 that these trends may reflect ongoing changes in stratification and the strength and depth of the 281 Atlantic Meridional Overturning Circulation (AMOC; e.g., Stouffer and Manabe 2003; Rugenstein 282 et al. 2016a; Marzocchi and Jansen 2017; Jansen et al. 2018). Even if the energy flux imbalance 283 at the TOA or the ocean surface are close to a new steady state this does not necessarily indicate 284 that the deep ocean is equilibrated as well (Zhang et al. 2013; Hobbs et al. 2016; Marzocchi and 285 Jansen 2017). Reaching deep ocean equilibration may not be necessary for studies concerned with 286

<sup>287</sup> surface properties only. However, for interpretation of paleo proxies and comparison with model
<sup>288</sup> simulations, distinguishing between the transient and equilibrium response in the intermediate or
<sup>289</sup> deep ocean is necessary (Zhang et al. 2013; Marzocchi and Jansen 2017; Rind et al. 2018; Jansen
<sup>290</sup> et al. 2018).

### <sup>291</sup> b. Evolution of surface temperature and cloud radiative effect

The evolution of large scale surface air temperature patterns on decadal to millennial time scales 292 (Fig. 5) are robust among models and different forcing levels. The simulations show a strong land-293 sea warming contrast on short time scales and little warming over the Southern Ocean on decadal 294 to centennial time scales (e.g., Manabe et al. 1991; Gregory 2000; Joshi and Gregory 2008; Geof-295 froy and Saint-Martin 2014; Armour et al. 2016). A warming pattern reminiscent of the positive 296 phase of ENSO and the Interdecadal Pacific Oscillation occurs throughout the Pacific basin (panel 297 b; Held et al. 2010; Song and Zhang 2014; Andrews et al. 2015; Luo et al. 2017) but decays on 298 centennial to millennial time scales (panel c and d), with a large model spread in time scales (not 299 shown). As it approaches equilibrium, the temperature pattern becomes more homogeneous, the 300 land-sea warming contrast reduces (e.g., Held et al. 2010; Geoffroy and Saint-Martin 2014), and 301 the Southern Hemisphere high latitudes keep warming beyond year 1000. As in previous studies, 302 the AMOC first declines (Gregory et al. 2005; Zhu et al. 2014; Kostov et al. 2014; Trossman et al. 303 2016) and then recovers (Stouffer and Manabe 2003; Li et al. 2013; Zickfeld et al. 2013; Rugen-304 stein et al. 2016a; Rind et al. 2018), resulting in a delayed warming in the North Atlantic. Panel 305 a, b, and e correspond to the blue shading in Fig. 1, and are known from CMIP5 simulations (e.g., 306 Andrews et al. 2015), while panel c, d, f, and g highlight that the simulations still warm substan-307 tially on centennial to millennial time scales, mainly in areas with more sensitive – i.e. positive 308 or small negative – feedbacks (Rugenstein et al. 2019). Normalizing the zonal-mean temperature 309

anomaly by the global mean warming reveals the relative zonal-mean warming (Fig. 6). Arctic am-310 plification begins very early in the simulations and warming throughout the Southern Hemisphere 311 is lower than the global average in almost all models for the first centuries. Between year 100 and 312 1000 the Southern Hemisphere warms more than the Northern Hemisphere in all latitudes pole-313 ward from 30°, in some regions by more than 4 K. Antarctic warming slowly increases, but is still 314 substantially less than Arctic amplification (e.g., Salzmann 2017). In a couple of models, the am-315 plitude of Antarctic and Arctic amplification is the same after 4000 years of model integration time 316 (GISSE2R and ECHAM5; Li et al. 2013), while in other models the Antarctic amplification stays 317 substantially smaller and still increasing after a couple of thousand years. LongRunMIP shows 318 that there is no reduction in model spread in the polar regions through time and that although all 319 models follow a similar large scale pattern evolution (Fig. 5), the local response time scales, e.g. 320 in the North Atlantic, Southern Ocean, or equatorial Pacific differ by hundreds to thousands years. 321 While the large scale temperature response is rather robust between models and simulations, 322 the cloud radiative effect (CRE) differs strongly in magnitude and time evolution, both between 323 models and between forcing levels for the same model (Fig 7). We show the shortwave CRE – 324 computed as the difference between "all sky" and "clear sky" shortwave radiative fluxes (e.g., 325 Ramanathan et al. 1989; Ceppi et al. 2017) – as a function of surface air temperature anomaly. 326 The models disagree in the overall sign, as expected from CMIP5 models on shorter time scales 327 (e.g., Vial et al. 2013; Caldwell et al. 2015), but can even change sign within a single simulation 328 (e.g., ECEARTH or CESM *abrupt8x*). The strength of variation in time within one simulation 329 can depend strongly on the forcing level (e.g. MIROC32 *lpct2x* vs. *lpct4x*) and the time scales 330 of change differ between the models (e.g. IPSLCM5R vs. MPIESM12 abrupt4x). For some 331 simulations, cloud response barely changes with temperature, contributing negligibly to the overall 332 feedback (e.g. MPIESM12 abrupt16x, CESM104 abrupt4x, and MIROC32 1pct2x). 333

### **4. Discussion and Outlook**

### *a. Published millennial-length simulations*

Models of intermediate complexity are the most common tools used to study century to millen-336 nium time scales in the climate system (e.g., Zickfeld et al. 2013; Eby et al. 2013; Levermann et al. 337 2013; Rugenstein et al. 2016c; Jansen et al. 2018). However, they usually have a poorly resolved 338 atmosphere and little or no representation of cloud processes. In contrast, the publications in Table 339 3 feature millennium-length AO-GCM simulations. Asterisks mark contributions to LongRunMIP. 340 These papers provide a solid body of work on millennial-length climate simulations, but rarely use 341 the same forcing levels and simulation length and focus on different aspects of the climate sys-342 tem. Three papers compare model formulation and processes of two AO-GCMs each (Frölicher 343 et al. 2014; Paynter et al. 2018; Krasting et al. 2018), but otherwise models have not been sys-344 tematically compared against each other. Fig. 4 and 7 show that AO-GCMs can strongly differ in 345 their behavior. Spatial patterns of e.g., precipitation and surface heat fluxes also vary strongly be-346 tween models and between different forcing scenarios for the same model (not shown), suggesting 347 that some mechanisms and processes discussed in the published literature are not generalizable 348 across models. For example, there is disagreement about which regions are thought to dominate 349 the changing feedback parameter (Senior and Mitchell 2000; Andrews et al. 2015; Meraner et al. 350 2013; Caballero and Huber 2013) or whether or not, and on which time scales, the AMOC recovers 351 from its initial reduction (Voss and Mikolajewicz 2001; Stouffer and Manabe 2003; Li et al. 2013; 352 Rind et al. 2018; Thomas and Fedorov 2019). Paleo climate simulations are often several thou-353 sand years long, however, they usually include boundary conditions such as ice sheets or changing 354 continental configurations, which differ from the ones used here. However, paleo climate studies 355 often discuss equilibration time scales and deep ocean temperature trends relevant to the types 356

of models included in LongRunMIP (e.g., Brandefelt and Otto-Bliesner 2009; Zhang et al. 2013;
 Klockmann et al. 2016; Marzocchi and Jansen 2017; Gottschalk et al. 2019).

### 359 b. Limitations

LongRunMIP analyses are currently limited mainly by the collected *variables* (Table 1). In-360 cluding cloud fields and 3D atmospheric temperature and humidity fields, for example, would 361 allow users to study atmospheric dynamics and radiative feedbacks in more detail. The *differ*-362 *ent forcing scenarios* of model contributions to LongRunMIP are both a strength and weakness. 363 Minimal requirements have encouraged a large number of contributions so far. However, study-364 ing a single forcing scenario requires model selection or scaling between different forcing levels. 365 Slab ocean simulations, which replace a model's dynamical ocean with a much shallower non-366 dynamical mixed-layer, are a computationally cheap tool to compare fast and slow time scales and 367 the relevance of surface warming patterns (Boer and Yu 2003c; Danabasoglu and Gent 2009; Li 368 et al. 2013). We hope to receive submissions of these simulations in the future, to allow analysis of 369 their utility. Century to millennial-time scales in the real world include more processes and *Earth* 370 System Feedbacks than are included in LongRunMIP simulations, such as the carbon cycle, vege-371 tation feedbacks, forcing agents other than  $CO_2$  (such as other greenhouse gases or aerosols), ice 372 sheets, glacial rebound effects, changes to continental configuration, and orbital variation. Further, 373 the real climate system is never in equilibrium or steady state, because the forcing continuously 374 changes (e.g., Köhler et al. 2017). These Earth system feedbacks and additional forcings must be 375 taken into account when comparing the LongRunMIP models with paleo proxies or when project-376 ing or predicting changes in future centuries or millennia. 377

### 378 c. Summary and expected impact

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LongRunMIP is the first archive of millennial-length simulations of complex climate models, 379 featuring 50 simulations of 15 models by 10 modeling centers under various forcing scenarios (Ta-380 ble 2). The archive provides an unprecedented opportunity to study the equilibrium response of a 381 large number of models to forcing. The variables included allow study of a range of phenomena 382 associated with the atmosphere, ocean, land, and sea ice (Table 1), and we expect LongRunMIP to 383 contribute to current discussions laid out in Section 1. This includes ocean heat uptake, sea level 384 rise, ocean circulation response to warming, large scale modes of variability, sea ice reduction, 385 polar amplification, precipitation variability, atmospheric dynamics, long-term memory in time 386 series, spatial warming patterns, ocean - atmosphere interactions, model spin-up techniques, the 387 relation of internal variability and forced response under different forcing levels, committed cli-388 mate response, and the relation of time and state dependence of fast feedbacks and Earth System 389 Feedbacks and processes. 390

LongRunMIP is a MIP of opportunity, without an argeed upon protocol, and is a result of the willingness of individual research groups to provide model output from simulations often conducted over years of real-world time. As a result, the experiments are not standardized, but most models provided a millennial-length simulation that begins with an abrupt quadrupling of CO<sub>2</sub> concentration. In addition to collecting simulations, we provide output with standardized formats and variable names, and include versions regridded to a common grid, as well as global averages. LongRunMIP builds upon a body of pioneering studies that looked at the behavior of models be-

<sup>399</sup> a diverse group of models that exhibit strikingly different behavior (Fig. 7), and hopefully encour-

yond the centennial scale (Table 3), LongRunMIP allows this sort of analysis to be applied across

<sup>400</sup> age others to look beyond the limitations and assumptions normally imposed by computational <sup>401</sup> constraints, to directly study the equilibration of the fully coupled atmosphere-ocean system.

### 402 Data access and sharing

LongRunMIP currently consists of 15 TB of data and available for download at https://data.iac.ethz.ch/longrunmip/. Fields shown in this paper can be accessed on https://data.iac.ethz.ch/longrunmip/BAMS/.

See www.longrunmip.org for more details on available variables, contact information, sample figures and videos, and links to join a discussion community. We will be collecting more simulations over the next couple of years.

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TABLE 1. Description of collected variables. 2D means spatial resolution of latitude and longitude, except for 824 msftmyz where it means latitude and depth. 3D means latitude, longitude, and depth. msftmyz is the sum of the 825 eularian, eddybolus, and submeso component. For so and thetao there are also February and September values 826 available for most models.

Shortname	Longname	Unit	Resolution
hfls	Surface Upward Latent Heat Flux	${\rm W}{\rm m}^{-2}$	monthly, 2D
hfss	Surface Upward Sensible Heat Flux	${ m W}{ m m}^{-2}$	monthly, 2D
pr	Precipitation on atmospheric grid	$kgm^{-2}s^{-1}$	monthly, 2D
psl	Sea Level Pressure	Pa	monthly, 2D
rlds	Surface Downwelling Longwave Radiation	${\rm W}{\rm m}^{-2}$	monthly, 2D
rlus	Surface Upwelling Longwave Radiation	${\rm W}{\rm m}^{-2}$	monthly, 2D
rlut	TOA Outgoing Longwave Radiation	${\rm W}{\rm m}^{-2}$	monthly, 2D
rlutcs	TOA Outgoing Clear-Sky Longwave Radiation	${ m W}{ m m}^{-2}$	monthly, 2D
rsds	Surface Downwelling Shortwave Radiation	${ m W}{ m m}^{-2}$	monthly, 2D
rsdt	TOA Incident Shortwave Radiation	${ m W}{ m m}^{-2}$	monthly, 2D
rsus	Surface Upwelling Shortwave Radiation	${\rm W}{\rm m}^{-2}$	monthly, 2D
rsut	TOA Outgoing Shortwave Radiation	${\rm W}{\rm m}^{-2}$	monthly, 2D
rsutcs	TOA Outgoing Clear-Sky Shortwave Radiation	${ m W}{ m m}^{-2}$	monthly, 2D
tas	Near-Surface Air Temperature	К	monthly, 2D
ts	Atmospheric surface temperature	К	monthly, 2D
sic	Sea Ice Area Fraction	%	monthly, 2D
msftmyz	Meridional Overturning Circulation	${ m m}^3~{ m s}^{-1}$	annual, 2D
tos	Sea surface temperature	К	annual, 2D
SOS	Sea surface salinity	psu	annual, 2D
wfo	Net water flux into sea water	$\rm kgm^{-2}s^{-1}$	annual, 2D
evs	Water evaporation	$kgm^{-2}s^{-1}$	annual, 2D
pr_ocn	Precipitation (rain and snow) on ocean grid	$\rm kgm^{-2}s^{-1}$	annual, 2D
tauuo	Surface downward wind stress in x direction	${\rm N}{\rm m}^{-2}$	annual, 2D
tauvo	Surface downward wind stress in y direction	${ m N}{ m m}^{-2}$	annual, 2D
so	Sea Water Salinity	psu	annual, 3D
thetao	Sea Water Potential Temperature	K	annual, 3D

TABLE 2. Overview of models and contributed simulations. The resolution of atmosphere and ocean is given in # of grid points per latitude x longitude, and latitude x longitude x depth, respectively. Models are referred to by their shortnames throughout the manuscript. Section 2b describes the forcing levels. References in the last column describe the models and simulations. Some simulations are published in their full length, some simulations contributed to LongRunMIP are the extensions of simulations discussed in the references, and some simulations are unpublished.

Model (shortname)	Forcing level shortname	Length (yrs)	Atmosphere resolution	Ocean resolution	Control sim (yrs)	Model and simulation
CCSM3 CCSM3	abrupt2x abrupt4x abrupt8x	3000 2120 1450	48 x 96	100 x 116 x 25	1530	Yeager et al. (2006) Danabasoglu and Gent (2009)
CCSM3 CCSM3II	abrupt2.4 abrupt4.8 lin2.4	3701 3132 3990	48 x 96	100 x 116 x 25	3805	Yeager et al. (2006) Castruccio et al. (2014)
CESM 1.0.4 CESM104	abrupt2x abrupt4x abrupt8x	2500 5900 5100	96 x 144	384 x 20 x 60	1320	Gent et al. (2011) Danabasoglu et al. (2012) Rugenstein et al. (2016c)
CNRM-CM6-1 CNRMCM61	abrupt2x abrupt4x	750 1850	128 x 256	180 x 360 x 75	2000	Voldoire et al. (2019) Saint-Martin et al. (2019)
EC-Earth-PISM ECEARTH	historical RCP8.5+	1270	160 x 320	292 x 362 x 42	508	Hazeleger et al. (2012) Svendsen et al. (2015)
ECHAM5/MPIOM ECHAM5	abrupt4x 1pct4x	1000 6080	48 x 96	101 x 120 x 40	100	Jungclaus et al. (2006) Li et al. (2013)
FAMOUS FAMOUS	abrupt2x abrupt4x	3000 3000	37 x 48	73 x 96 x 20	3000	Smith et al. (2008)
GFDL-CM3 GFDLCM3	1pct2x	5000	90 x 144	200 x 360 x 50	5200	Donner et al. (2011) Paynter et al. (2018)
GFDL-ESM2M GFDLESM2M	1pct2x	4500	90 x 144	200 x 360 x 50	1340	Dunne et al. (2012) Paynter et al. (2018)
GISS-E2-R GISSE2R	abrupt4x	5000	90 x 144	180 x 288 x 32	5225	Schmidt et al. (2014); Miller et al. (2014); Nazarenko et al. (2015)
HadCM3L HadCM3L	abrupt2x abrupt4x abrupt6x abrupt8x	1000 1000 1000 1000	73 x 96	73 x 96 x 20	1000	Cox et al. (2016) Cao et al. (2016)
HadGEM2-ES HadGEM2	abrupt4x	1328	145 x 192	216 x 360 x 40	239	Collins et al. (2011) Andrews et al. (2015)
IPSL-CM5A-LR IPSLCM5ALR	abrupt4x	1000	96 x 96	149 x 182 x 31	1000	Dufresne et al. (2013)
MIROC 3.2 MIROC32	1pct2x 1pct4x	2000 2000	64 x 128	192 x 256 x 44	681	Hasumi and Emori (2004) Yamamoto et al. (2015); Yoshimori et al. (2016)
MPIESM-1.2 MPIESM12	abrupt2x abrupt4x abrupt8x abrupt16x	1000 1000 1000 1000	96 x 192	220 x 256 x 40	1237	Mauritsen et al. (2018) Rohrschneider et al. (2019)
MPIESM-1.1 MPIESM11	abrupt4x	4459	96 x 192	220 x 256 x 40	2000	Mauritsen et al. (2018)

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Paper	Model	Forcing level	Length (yr)	Content/scientific comment
Senior and Mitchell (2000)	HadCM2	2xCO <sub>2</sub>	pprox 800	Included flux adjustments; effective climate sensitivity increases due to SW CRE due to changes in the inter-hemispheric temperature gradient
Bi et al. (2001)	CSIRO	3xCO <sub>2</sub>	≈ 1000	Cessation and recovery of Antarctic Bottom Water and North Atlantic Deep Water forma- tion
Voss and Mikolajewicz (2001)	ECHAM3	2x and 4xCO <sub>2</sub>	850	Adjustment time scales, committed warming, ocean thermohaline circulation
Stouffer and Manabe (1999 2003)	, GFDL	0.5x, 2x, and 4xCO <sub>2</sub>	1 4000	Thermohaline circulation and paleo-oceanographic implications
Boer and Yu (2003b,a,c)	CCCma	2050 and 2100 forcing	0 1000	Radiative feedbacks and surface warming; effective climate sensitivity decreases with time; slab versus fully coupled models
Gregory et al. (2004)	HadCM3	2xCO <sub>2</sub>	$\approx 1000$	TOA radiative imbalance and surface temperature are not linearly related; after 1000 yr the model is still $0.7$ W m $^{-2}$ away from equilibrium
* Danabasoglu and Gen (2009)	t CCSM3	2x, 4x, and 8xCO <sub>2</sub>	3000	Comparing slab and fully coupled models; determining ECS; Jonko et al. (2013) analyzed the contributions of different feedbacks to doublings of CO <sub>2</sub>
Gillett et al. (2011)	CanESM1	21st century	$\approx 1000$	Impact of reduced emissions
* Li et al. (2013)	ECHAM5/MPI-OM	2xCO <sub>2</sub>	≈ 6000	Comparing slab and fully coupled models; determining ECS; adjustment time scales of surface warming patterns, ocean heat uptake, and sea level rise
Frölicher et al. (2014) Frölicher and Paynter (2015)	; GFDL-ESM2M, CSM1	4xCO <sub>2</sub> pulse	1000	Climate impact of CO <sub>2</sub> emission stoppage; evolving feedbacks; ECS; transient climate response to cumulative carbon emissions
* Andrews et al. (2015)	HadGEM2-ES	4xCO <sub>2</sub>	$\approx 1300$	Non-constancy of feedbacks; variations of TOA components cancel each other on the century to millennial time scale
* Yamamoto et al. (2015) Yoshimori et al. (2016)	; MIROC 3.2	$2x$ and $4xCO_2$	2000	Deep ocean ventilation overall increases oxygenation after a transient decrease; review article on ocean heat uptake in coupled models and energy balance models
* Cao et al. (2016)	HadCM3L	2x, 4x, 6x, 8xCO <sub>2</sub>	1000	Comparing CO <sub>2</sub> to other forcing agents and geo-engeneering scenarios
* Rugenstein et al. (2016b,a)	CESM104	2x, 4x, 8xCO <sub>2</sub>	$\approx 1300$	Dependence of global and regional radiative feedback evolution on surface heat flux pat- ters; forcing adjustment
* Paynter et al. (2018)	GFDL-ESM2M, GFDL- CM3	- 2xCO <sub>2</sub>	$\approx 5000$	Evolution of global and regional radiative feedbacks and the role of atmospheric vertical velocity fields and inversion strengths
* Rind et al. (2018)	GISS-E2-R	4xCO <sub>2</sub>	$\approx 2000$	AMOC reduction and recovery on North Atlantic surface flux conditions
Krasting et al. (2018)	GFDL-ESM2Mb, GFDL-ESM2G	4xCO <sub>2</sub>	5000	Ocean heat uptake, model formulation of diapycnal diffusivity and ocean vertical coordinates

# TABLE 3. Published millennial-length simulations

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FIG. 1. Global and annual mean surface air temperature (*tas* in Table 1) anomaly and top of the atmosphere (TOA) radiative imbalance (computed as *rsdt* - *rlut* - *rsut*, see Table 1) to a step-forcing of quadrupling  $CO_2$ as simulated by the CESM104 model. For the Coupled Model Intercomparison Project Phase 5 and 6, this simulation is part of the standard protocol, but only 150 simulated years are requested (blue shading). We collect simulations that extended this experiment for at least 850 years (light red shading), ideally until they are equilibrated (end of dark red shading).



FIG. 2. Global annual mean surface air temperature for all control (black) and forced (color, listed in the top right of each panel) simulations. *abrupt2x*, *4x*, *6x*, *8x* means that the CO<sub>2</sub> concentration is doubled, quadrupled, sextupled, octupcliated, as a step-forcing branched off the control simulation. *1pct2x* and *1pct4x* means the CO<sub>2</sub> concentration is linearly increased 1 % per year until the concentration is doubled or quadrupled, respectively. The simulations of ECEARTH and CCSM3II are described in Section b. Note the different axis ranges for each model. GFDLCM3 and CCSM3II are not branched off directly from the control simulation.



FIG. 3. Top of the atmosphere (TOA) annual and global mean radiative imbalance of all control simulations.
 Note the different lengths of the horizontal axes. The gray line indicates the average, the red line the linear trend,
 except for CCSM3II and GFDLCM3 for which both colors depict a fourth-order-polynomial fit.



FIG. 4. Global and annual mean temperature anomalies (experiment minus average of the control simulation) of the surface ocean (a, first layer) and deep ocean (b), as well as absolute values of deep ocean temperature in the control simulations (c), for *abrup4x* (solid) and *lpct4x* (dashed) simulations. "Deep ocean" means around 2000 m depth (closest level). Note that the time scale in c) is shorter than in a) and b).

a) Temperature anomaly year 15-25



b) Temperature anomaly year 80-120



c) Temperature anomaly year 900-1000



2 3 4 5 6 7 8 9 10 11 12 K

e) Change in temperature anomalies (b)-(a)



f) Change in temperature anomalies (c)-(b)



FIG. 5. Time evolution of the surface air temperature anomaly in the *abrupt4x* simulations. The model mean of CCSM3, CESM104, CNRMCM61, ECHAM5, GISSE2R, HadCM3L, HadGEM2, IPSLCM5A, MPIESM11, and MPIESM12 is shown in panel a, b, c, e, and f, while the model mean of only CESM104, GISSE2R, and MPIESM11 is shown in panel d and g, due to the length of these contributions. See Table 2 for details of the length of each simulation.



FIG. 6. Time evolution of the zonal mean surface air temperature response normalized by the global mean temperature anomaly. Above (below) 1 means that warming is amplified (reduced) relative to the globally mean warming (a-d). Panel e-g show the differences (note the difference scale). Panel a, b, e, and f contain only *abrupt4x* simulations, while panel c, d, and g also contain the *1pct2x* and *RCP8.5*+ simulations with integration lengths above 4000 years. Table 2 lists all simulations and model long names.



FIG. 7. Simulated shortwave cloud radiative effects SW CRE for different levels of global surface air temperature changes. Each point is a ten-year running average. Note the different axes labels, which cover a large range in TOA imbalance and surface temperature. Table 2 lists all simulations and model long names.